



### Common Astronomy Software Applications

## CASA Parallelization and High Performance Computing

Concepts, Implementation, and Issues

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## **CASA** and **HPC**

- CASA does not have a HPC issue.
  - We must support efficient operation on the available hardware.
- CASA has focused on two "standard" systems:

### **Workstation**

- Multi-core system
- Local disk
- Single shared memory

### <u>Cluster</u>

- Many multi-core nodes
- High performance network file system (Lustre)
- No shared memory access



# CASA

## Definitions

- Core: A single processing element which reads and executes instructions.
- Node: A single host, in modern systems usually has multiple cores.
- Engine: A single instance of CASA connected through a messaging interface, in most cases a single engine consumes a single core.
- Master Engine: The primary engine which is responsible for dispatching jobs to other engines
- Multi-Process: Many independent processes each with their own process space.
- Muti-Threaded: A single process with multiple threads of execution, multiple threads can share the same memory object.



# CASA The processing performance problem

• Many tasks in CASA require traversing the entire data set and are data IO limited.

		OS Realized Data Rate	Time for 500 GB file
SATA Disk	115 MB/s	60 MB/s	2h 22 m
Raid	200-500 MB/s	375 MB/s	22 m
Lustre (10 GB)	I.2 GB/s	700 MB/s	12 m

• Even imaging [in the simplest case] is IO limited, requiring about 50 MB/s per core to prevent the CPU from incurring wait states.

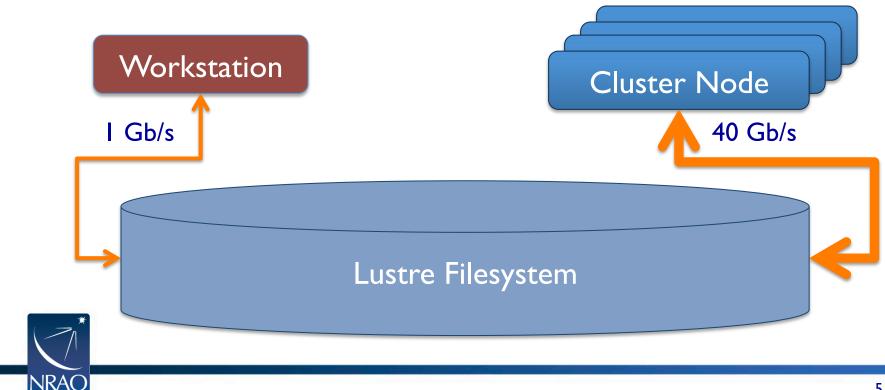
https://science.nrao.edu/facilities/evla/data-processing/hardware-recommendations





### Lustre

- Running against Lustre can provide significant performance improvement, ۲ but where you run from matters.
- How the data is arranged on Lustre impacts performance as well. ٠



# CASA Embarrassingly Parallel Approach

An embarrassingly parallel workload is one for which little or no effort is required to separate the problem into a number of parallel tasks, this is often the case where there is exists no dependency between the parallel tasks.

-Wikipedia

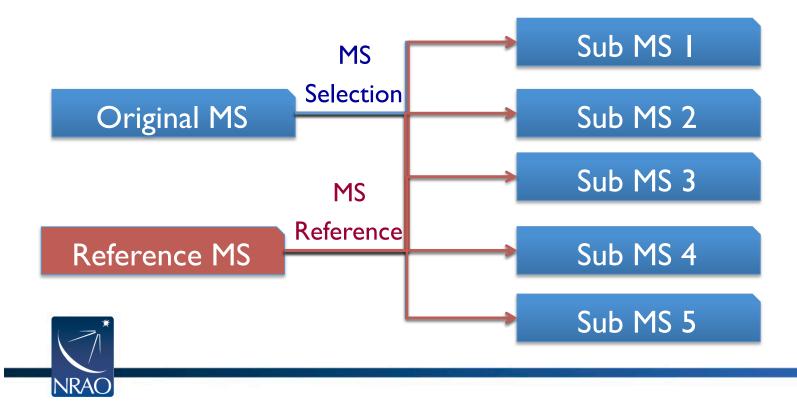
- Most of the tasks which require access to large amounts of data fit this description.
  - Flagging of Data
     Time Averaging Data
  - Applying Calibration
- Imaging does not strictly fit this definition but can the communication between processes is only at certain points in the cycle so it is also an <u>easily solved issue</u>.





### **Reference MSs**

- The easiest way to parallelize in CASA is to have multiple instances of CASA each running on a subset of the data.
  - To simplify the interface and minimize bookkeeping for the user we use a reference MS.





### **Partition**

- To simplify creation of the reference MS, use the task Partition.
- Contains usual data selection parameters:
  - Data column

• Antenna

• Array

UV-Range

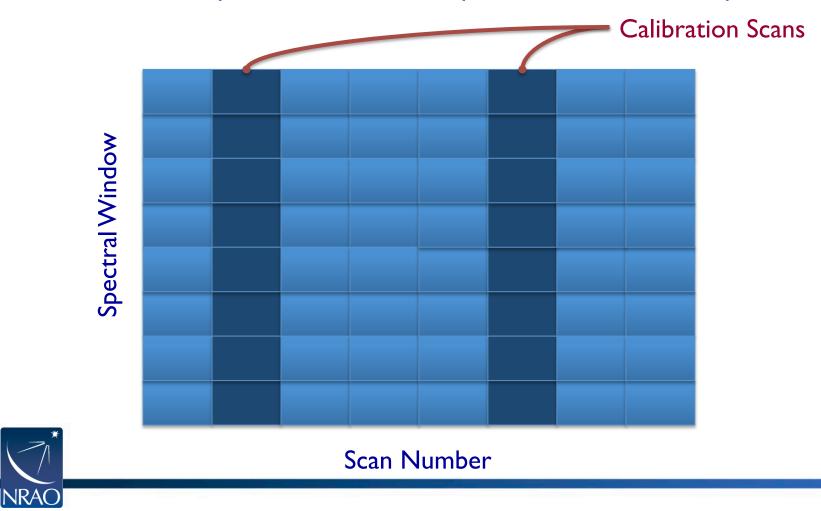
- Field
- Spw

- Scan
- Scan Intent
- If createmms is true a reference MS is created which contains at least numsubms sub MSs
- Partitioning also has the effect of distributing the sub-MSs across multiple raid controllers on Lustre, thus improving I/O performance.



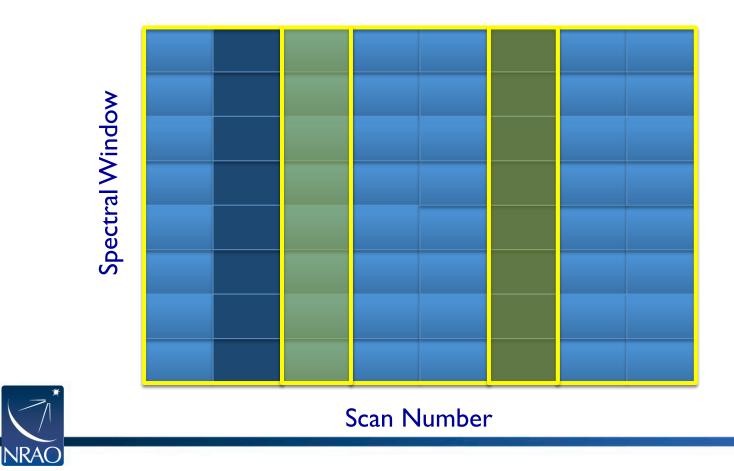


• Partition accepts three axis to do separation across default, spw, scan



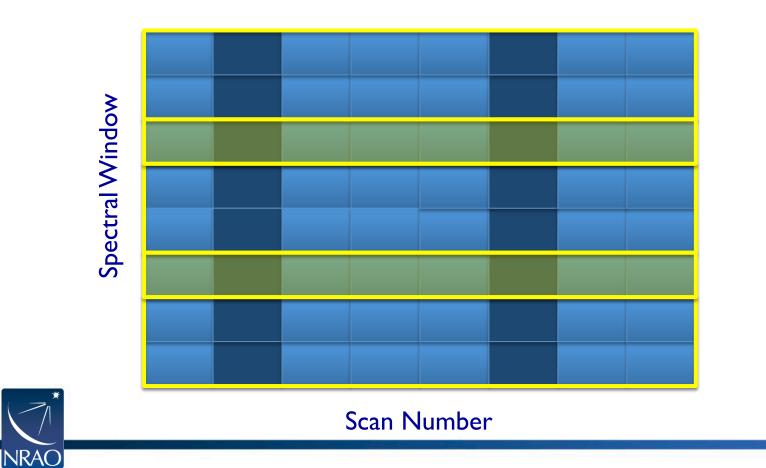


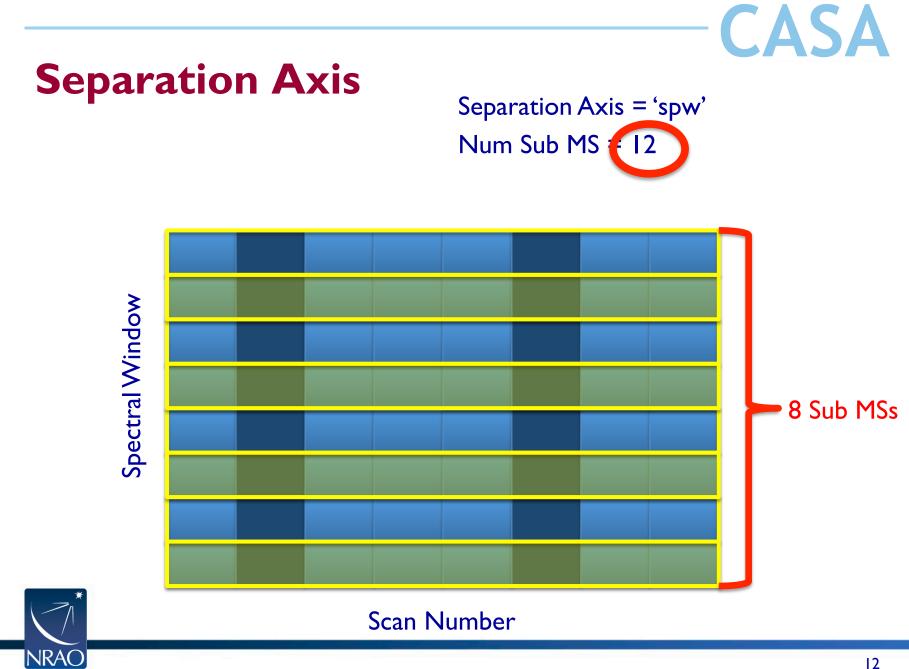
Separation Axis = 'scan' Num Sub MS = 5





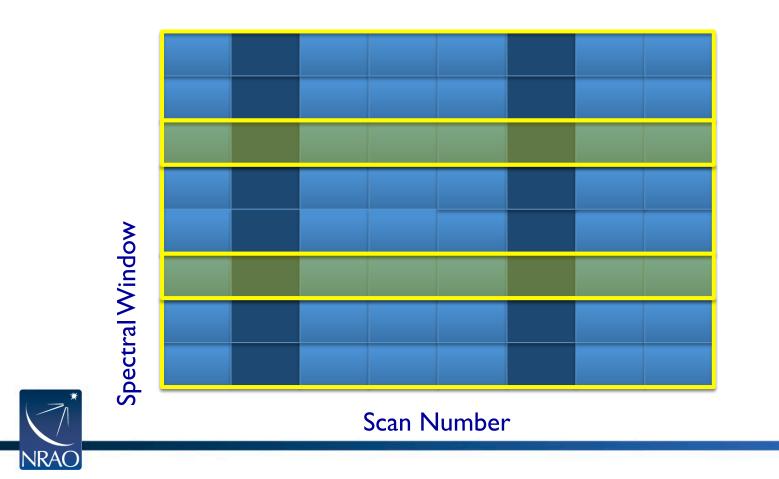
Separation Axis = 'spw' Num Sub MS = 5





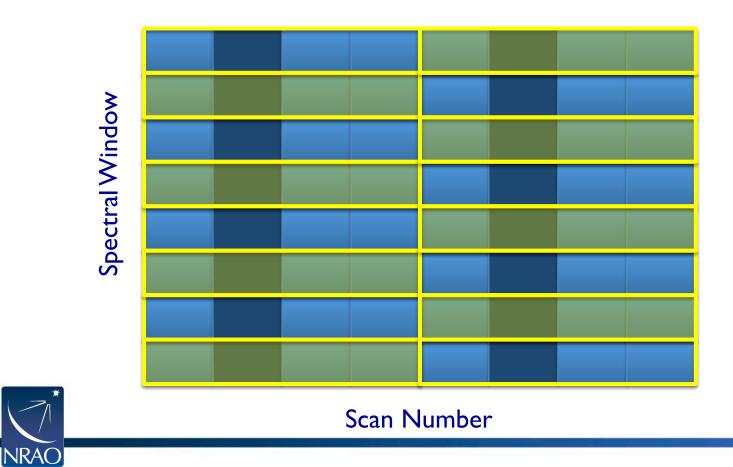


Separation Axis = 'default' Num Sub MS = 5



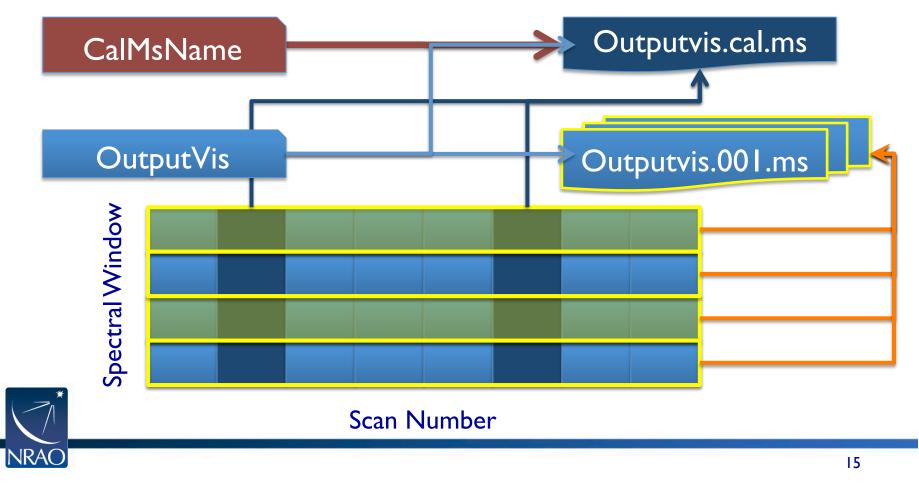


Separation Axis = 'default' Num Sub MS = 12





- Setting calmsselection to 'manual' allows selection on Field, Scan, or Intent.
- 'auto' does a selection based on scan intent





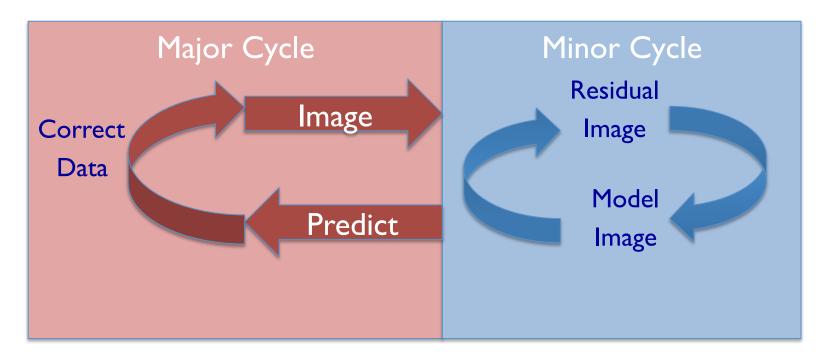
## **Pre-Imaging Tasks**

- After partitioning, any task which works on the MS in place (e.g. does not create a new MS) and does not modify the subtables can be parallelized fairly easily.
- So far I have only done those tasks which offered the most significant performance improvement for minimal investment :
  - ClearCal FlagData ApplyCal
- More tasks can easily be added as we identify which are most necessary
  - Flag\_cmd
  - SetJy



# CASA

## Imaging



• Parallelization of imaging does not depend on how the data is partitioned in any way.

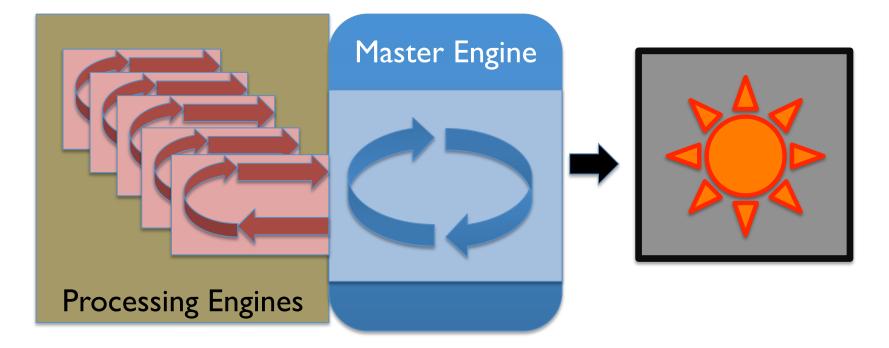


 Partitioning is a simple way to distribute the data more efficiently within the Lustre file system.



## **Continuum Imaging**

• So far for continuum imaging we have only parallelized the major cycle.







• So far parallel clean is only available at the tool level.

>>> from parallel.pimager import pimager

>>> imager =pimager()

>>> imager.pcont(arguments)

• We plan to incorporate this into the task once a few more features have been enabled and after we have sufficient experience with users to be confident.





- Controlling the switch between major and minor cycles and how deeply to clean is fairly manual at the moment.
- The user specifies:
  - majorcycles integer number of CS major cycles to do
  - niter: maximum number of clean iterations to execute
  - threshold: string quantity of the residual peak at which to stop deconvolving
    - e.g. '0.1 mJy'
- Between each CS major cycles the minor cycle will fit niter/majorcycles components.
- If threshold is reached at any time the clean cycle will stop.





- contclean: Boolean specifying if we are continuing a clean or starting new
  - If false the imagename.model is deleted otherwise clean will continue from previous run

By setting contclean = True the clean process can be resumed to clean deeper or with a new mask.





- The online help has definitions for the arguments many are the same as normal imaging:
- msname: input measurement set
- imagename: output image
- imsize: list of 2 numbers [nx,ny] defining image size in x and y
- pixsize: list of 2 quantities ['sizex', 'sizey'] defining the pixel size
  - e.g ['larcsec', 'larcsec']
- phasecenter: an integer (field index) or a direction
  - e.g 'J2000 19h30m00 -30d00m00'
- field: field selection string
- spw: spw selection string





- stokes: string specifing imaging stokes
  - e.g 'l', 'IV
- ftmachine: the ftmachine to use (see Sanjay's lecture)
  - ft, wproject, mosaic NOTE: wproject and mosaic are still being tested
- facets: integer the number of facets to split image into.
- alg: string specifying the algorithm
  - Current possibilities are 'clark', 'hogbom', 'msclean'
- weight: string
  - e.g 'natural', 'briggs' or 'radial'
- robust: float valid for 'briggs'
- scales: scales to use when using alg='msclean'





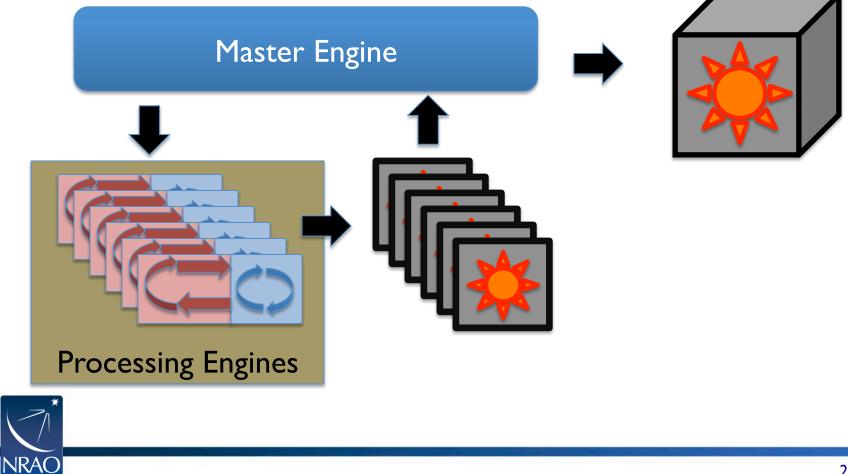
- These parameters are either not yet fully implemented, have been deprecated, or are available for developer use.
  - hostnames: deprecated use " (empty string)
  - numcpuperhost: deprecated use any integer
- visinmem: Boolean load visibility in memory for major cycles set to False
- These parameters are to support the work ongoing work on A-Projection, simply use the default values.
  - painc, cfcache, pblimit, dopbcorr, applyoffsets epitablename





## **Spectral Line Imaging**

• Each plane of the output cube can be done independently.



### **Running Parallel Spectral Imaging**

>>> from parallel.pimager import pimager

>>> imager =pimager()

>>> imager.pcube(arguments)

- Like parallel continum clean there is no task (yet) for parallel spectral line imaging.
- Most arguments are the same as for the Continuum case
- The *maskimage* parameter is an input image mask. If it is not the same shape as the final cube a best guess interpolation is done.
- Interactive clean is not supported in this mode, suggestions are welcome.



CASA



## **Running Spectral Line Imaging**

- A new parameter has been introduced in the *pcube* method: *chanchunk* 
  - chanchunk determines how many output channels are handled simultaneously by each of the processing engines.
  - The trade off here is between the amount of I/O and the memory footprint.

#### Serial Reduction: 351 min

Once we have understood the heuristics we expect to handle this parameter in the Task.

	Number of Engines		
Chan Chunk	6	H	12
I	123 min	98 min	96 min
4	73 min	51 min	
8	59 min	43 min	
16	57min		





### **Resource Specification**

- The computing hardware resources available to CASA are specified in a cluster specification file.
  - By default CASA assumes that it should use all of the cores available on the current system.
- Specification file looks like:

# <hostname>, <number of engines>, <working dir> casa-dev-06, 4, /lustre/jkern/parallelTest/cont.large casa-dev-07, 8, /lustre/jkern/parallelTest/cont.large

- You must be able to ssh without typing a password to all of the systems listed in your specification file.
  - If you do not have a specification file you must be able to ssh back to local host



See the NRAO goldbook for instructions on setting up ssh.



### **Resource Specification**

In order to instruct CASA to use the resources in your cluster ۲ configuration file:

>>> from simple\_cluster import simple\_cluster
>>> simple\_cluster().init\_cluster(<clusterFile>)

After these lines all tasks or tools which can make use of the ٠ parallelization will use these resources.





## **Advanced Usage**

- For experts there are some helper functions to assist in making use of the parallelization.
- JobData is a python class which encapsulates work to be done on a single engine:

>>> from simple\_cluster import JobData

>>> myJob = JobData('flagdata', <Argument Dictionary>)

• JobQueueManager is responsible for executing a set of jobs on the cluster

>>> from simple\_cluster import JobQueueManager
>>> queue = JobQueueManager()
>>> queue.addJob(myJob)
>>> queue.executeQueue()

• This will return when all jobs have either succeeded or broken. Return state can be found from the queue object.



Check the online help for more details on using these classes.