

# System Considerations

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# Considerations

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- Cost effectiveness
  - trade-off between specifications and cost
  - minimize operations cost
- Reliability
  - minimize mechanical devices
  - component lifetime calculations
  - extensive monitoring (and automatic diagnostics)
  - maintainability

# Precedents

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- JVLA
  - similar frequency range
  - older technology
- ALMA
  - overlap in frequency
  - new technology
  - similar cost

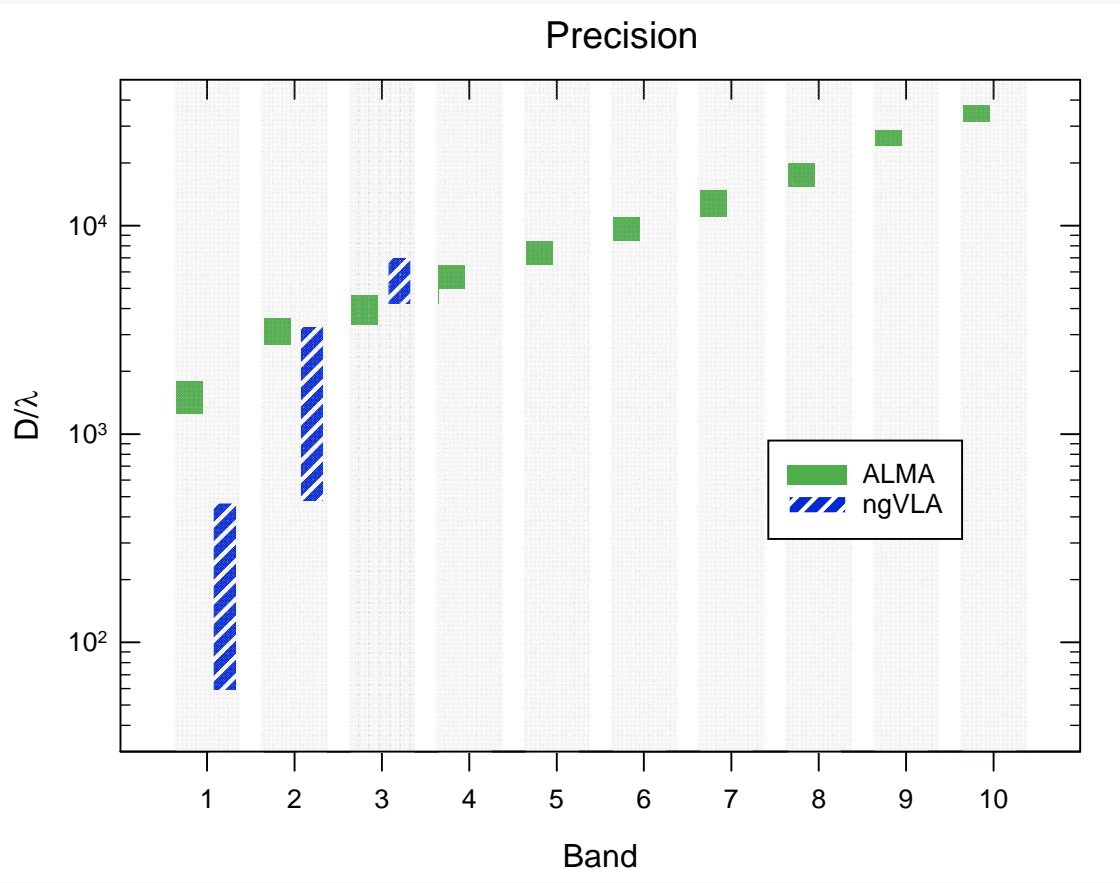


Image courtesy of NRAO/AUI

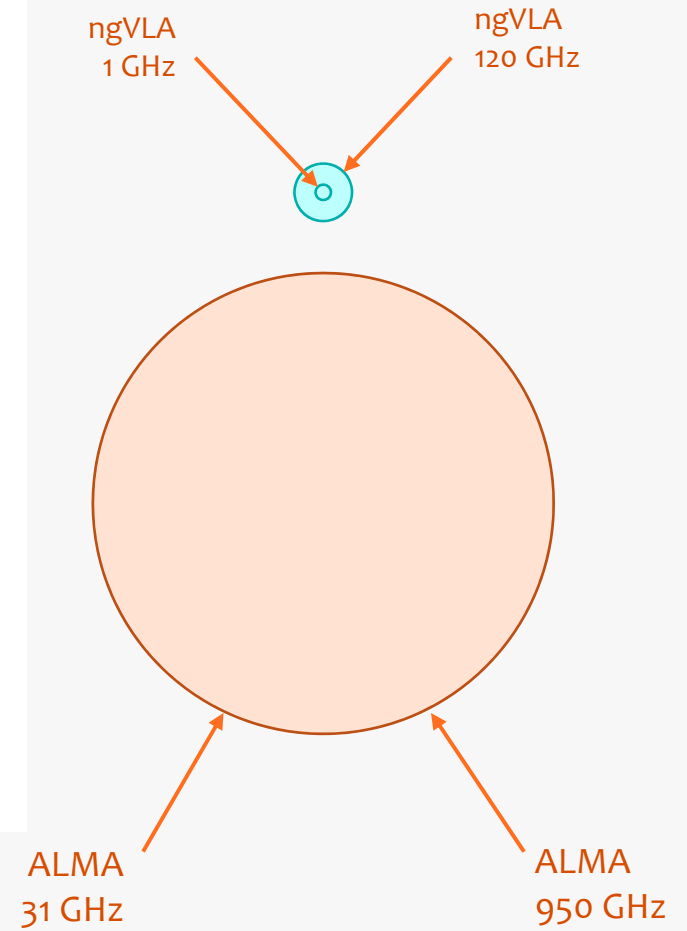


Image courtesy of NRAO/AUI and ALMA/ESO/NRAO/NAOJ

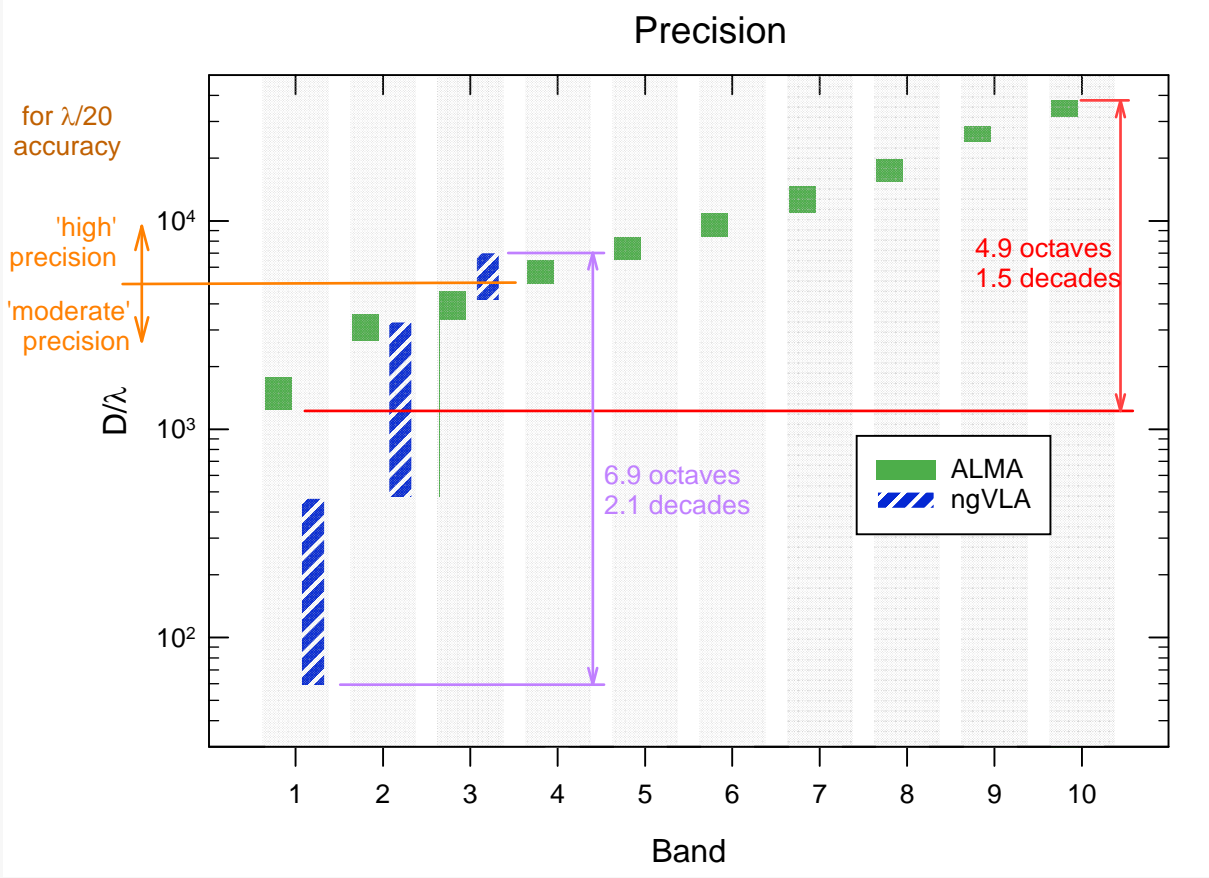
# Comparison of ALMA and ngVLA



antenna size in wavelengths



# Some observations



- Considerable overlap
- Mostly, ngVLA is does not have high precision requirements
- Range of resolution is >100
- Keck: 1.2 decades, 4.1 octaves

# System design issues

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- offset or symmetric
- prime focus or Cassegrain/Gregorian
- phase correction ( $\lambda$  3-mm, long baselines)
- analog vs digital transmission
- fiber links
- thermal control
- maintenance
- ...

# Mechanical

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- Motors
  - antenna drive motors
  - cryogenic refrigerator
  - air conditioning
  - fans
  - focus (x, y, z)
  - receiver select

# Metrics?

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- G/T is a common metric for antennas
- for ALMA
  - assumed this was  $\approx$  aperture efficiency
  - for electrically small antennas need *G and T*
- but is this the best metric?
  - yes, for point sources ( $\theta_{\text{source}} < \theta_{\text{FWHM}}$ )
  - for extended sources?
    - beam efficiency?
    - photons on sky?
- Other metrics
  - polarization purity
  - ...



# Aberrations for offset feed

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- Reduction in peak gain

Astigmatism: 
$$\Delta\eta_{ast} = \left( \frac{\pi D^3}{16Mfd\lambda} \right)^2 \alpha^4 \quad (1)$$

Coma: 
$$\Delta\eta_{com} = \frac{1}{2} \left( \frac{\pi D^3}{96M^2 f^2 \lambda} \right)^2 \alpha^2 \quad (2)$$

Curvature: 
$$\Delta\eta_{cur} = \frac{1}{3} \left( \frac{\pi D^3}{16fd\lambda} \right)^2 \alpha^4 \quad (3)$$

# Prime focus vs Cassegrain/Gregorian

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	prime focus	secondary focus
focal length	limited choice 0.25—0.8	free parameter 0.5—20
feed spillover	on ground	on sky
rx location	at prime focus	at secondary focus
field of view	small	large

# Feed sizes

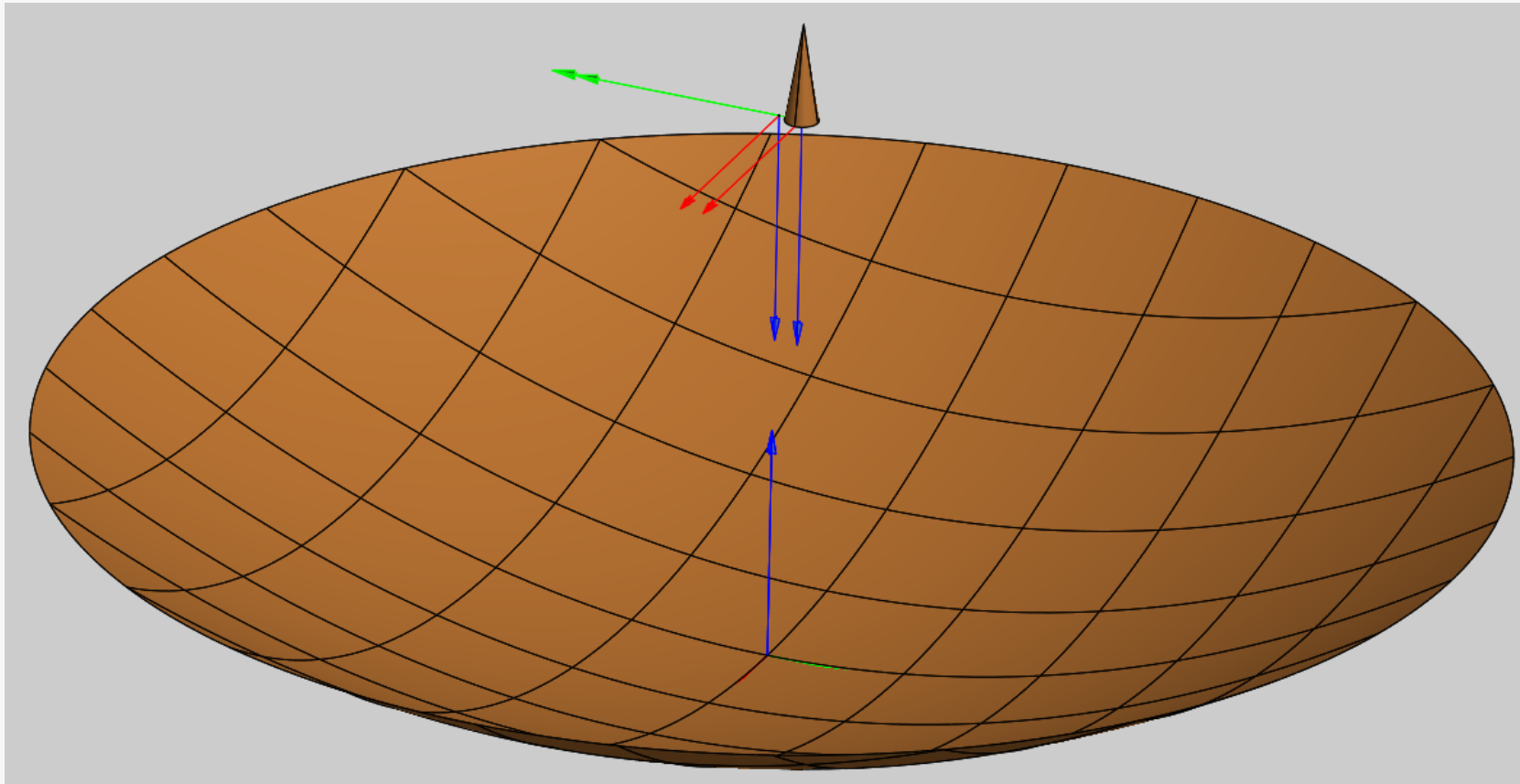
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feed diameter:  $d_f \approx 2.4 \lambda f/D$

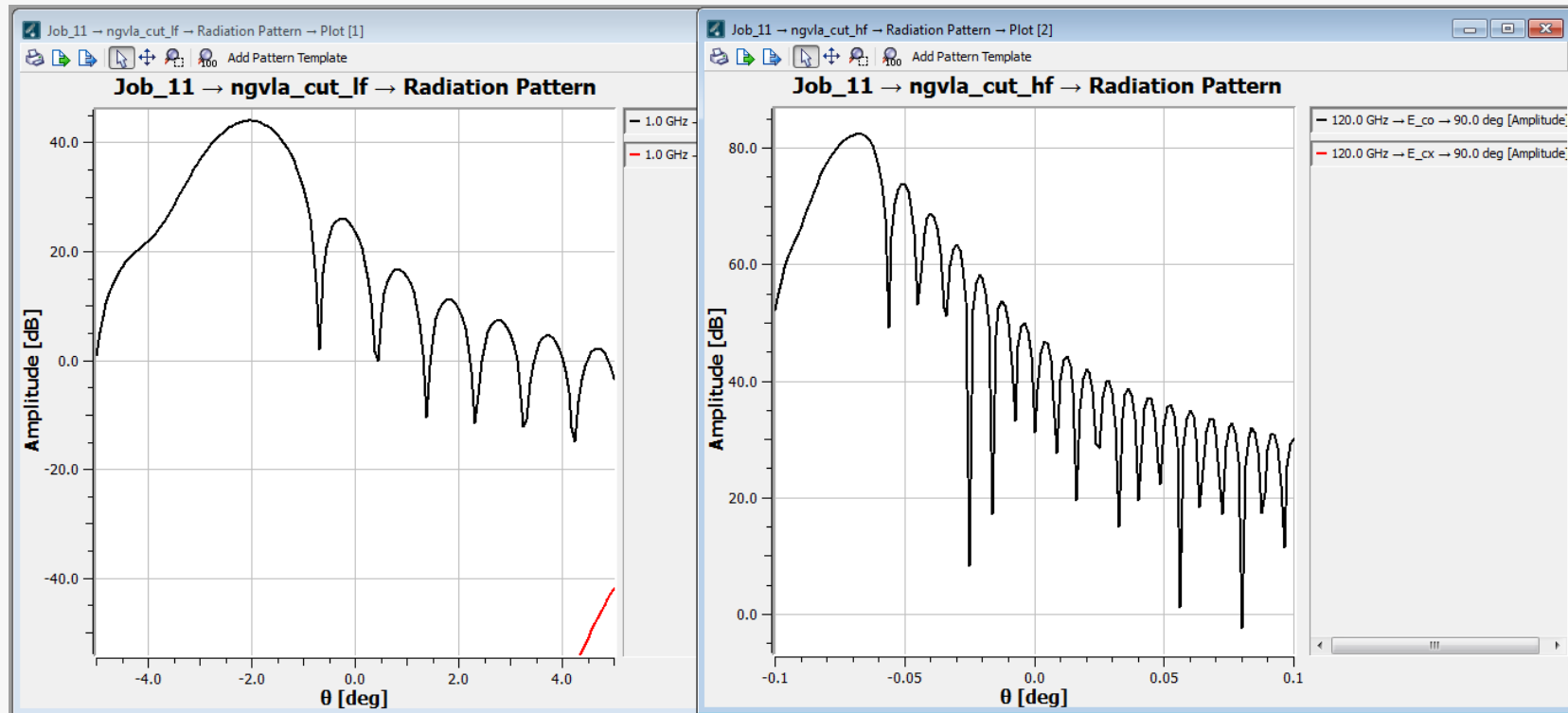
	prime focus	secondary focus	
f/D	0.4	0.6	5
1 GHz	290 mm	430 mm	3600 mm
8 GHz	36 mm	54 mm	450 mm
70 GHz	4 mm	6 mm	51 mm

# Offset feeds

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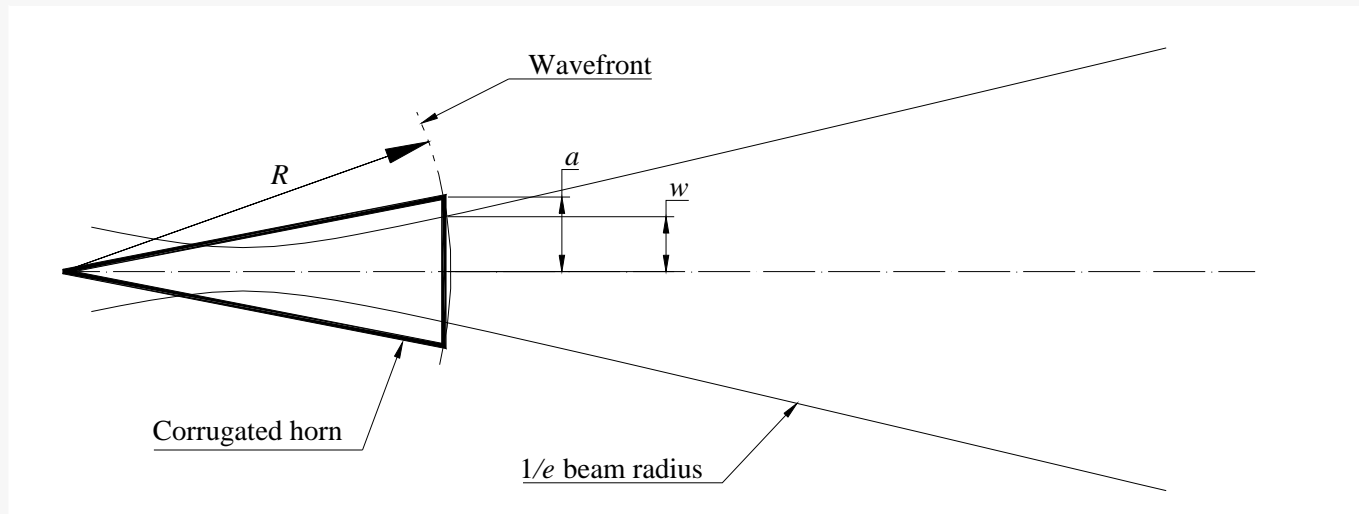


# Beam patterns



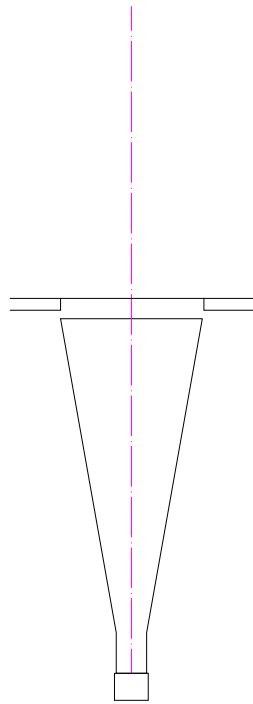
# Feed bandwidth

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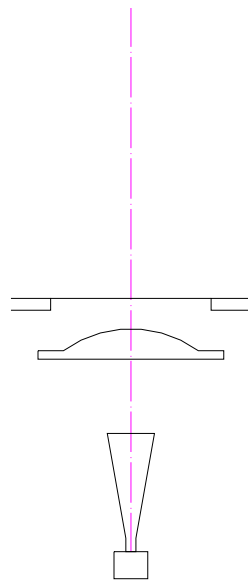


# Feed options

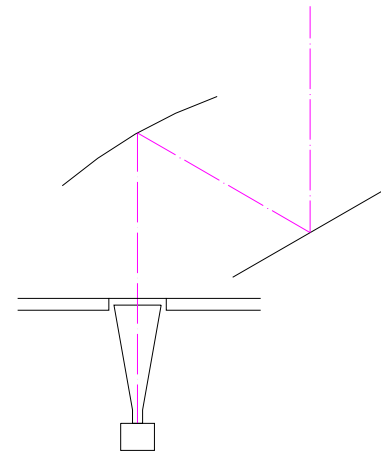
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horn only

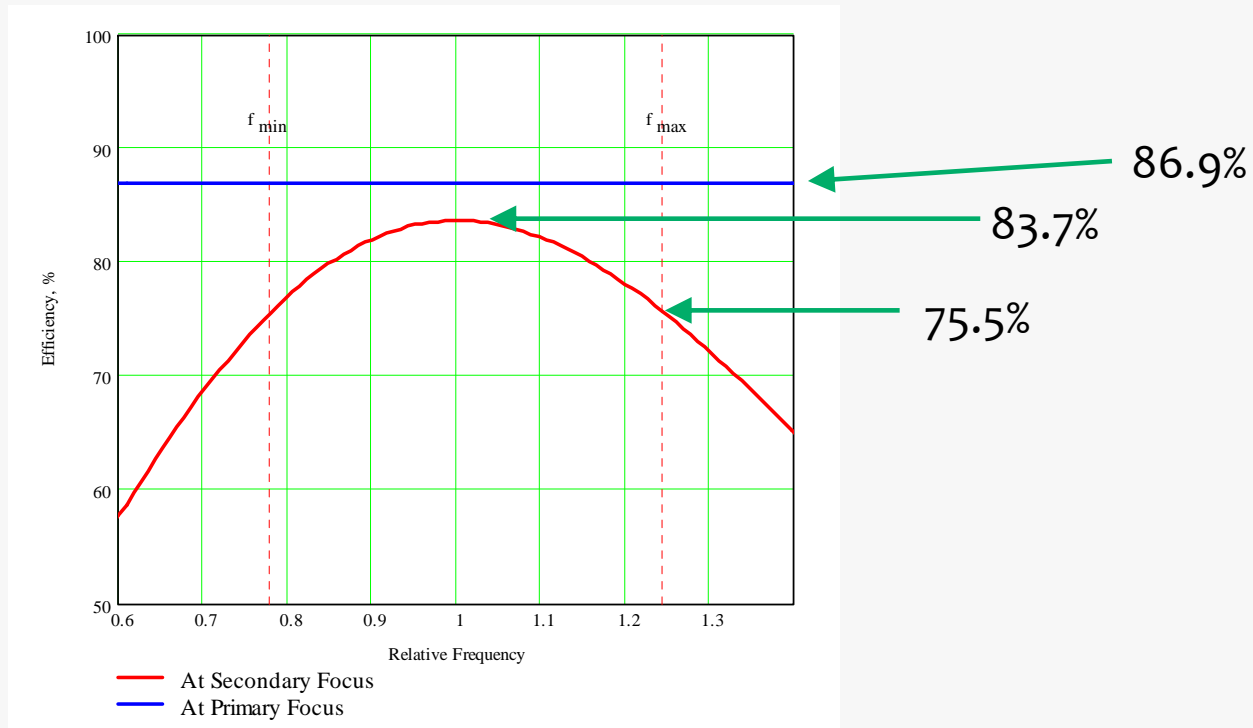


horn + lens



horn + offset ellipsoid + flat

# Corrugated horn (gold standard)

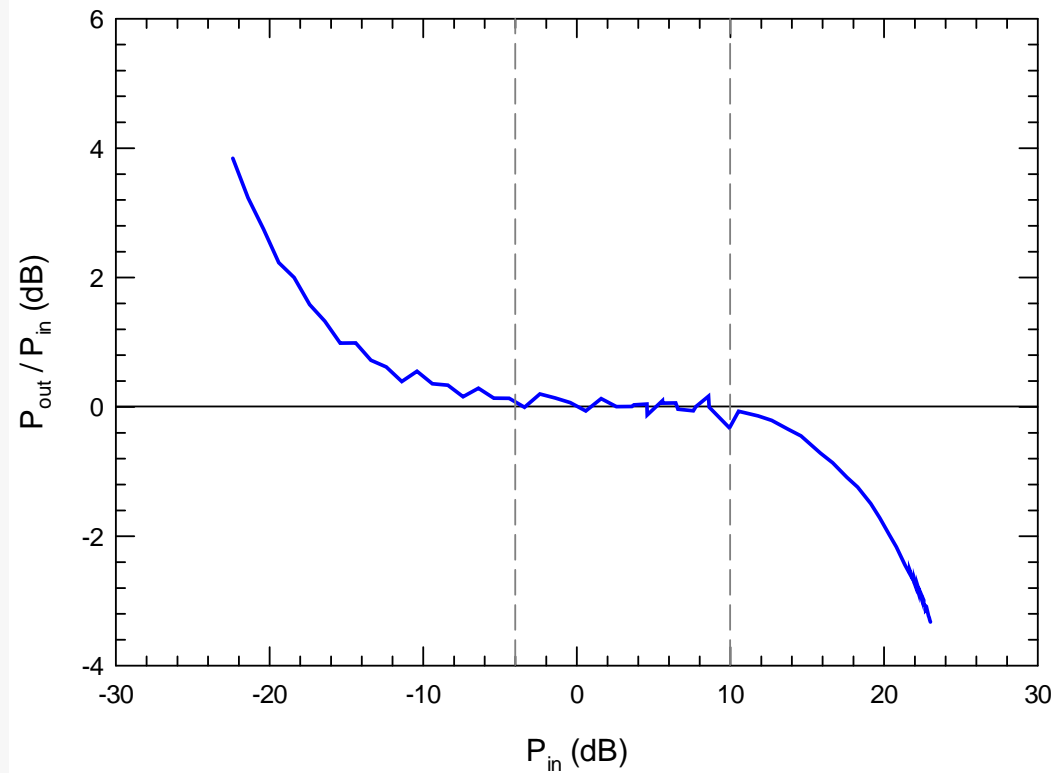


← 40% BW →  
 (need ~55% for ngVLA)



# Analog fiber link dynamic range

## Optilab fiber link



- input BW  $\sim 18$  GHz
- DR  $\sim 14$  dB
- pre-emphasized for uniform SNR
- DR scales  $\sim 1/BW$
- DR  $\sim P_{optical}$

# Conclusions

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- To early for conclusions
- Science drivers have to be clearly articulated
- Need a clear view on how much we want to trade off performance for cost