Computing for ngVLA: Lessons from LSST

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A Dedicated Survey Telescope



- A wide (half the sky), deep (24.5/27.5 mag), fast (image the sky once every 3 days) survey telescope. Beginning in 2022, it will repeatedly image the sky for 10 years.
- The LSST is an integrated survey system. The Observatory, Telescope, Camera and Data Management system are all built to support the LSST survey. There's no PI mode, proposals, or time.
- The ultimate deliverable of LSST is not the telescope, nor the instruments; it is the <u>fully reduced data</u>.
 - All science will be come from survey catalogs and images

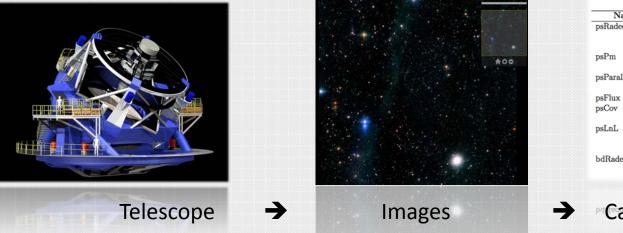


Table 4:	Level	2	Catalog	Object	Table
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Name	There a	Unit	Deceriation
	Type		Description
psRadecTai	double	time	Point source model: Time
			at which the object was at
			position radec.
psPm	float[2]	mas/yr	Point source model: Proper
-		,.	motion vector.
psParallax	float	mas	Point source model: Paral-
			lax.
psFlux	float[ugrizy]	nmgy	Point source model fluxes ⁵⁸ .
psCov	float[66]	various	Point-source model covari-
pacov	noar[00]	various	ance matrix ⁵⁹ .
	a .		direc mount i
psLnL	float		Natural log likelihood of
			the observed data given the
			point source model.
bdRadec	double[2]	degrees	B+D model ⁶⁰ : (α, δ) posi-
			tion of the object at time
			radecTai, in each band.
			radecTai. in each band.
			tion of the object at time
Catalo	in a second	degrees	B+D model ^{oo} : (α, δ) posi-
Catalo	Jes	dogroos	
	0-		point source model.
opsLnDeroorgeorgeorge	float		Natural log likelihood of

Open Data, Open Source: A Community Resource



- LSST data, including images and catalogs, will be available with <u>no</u> <u>proprietary period</u> to the astronomical community of the <u>United States</u>, <u>Chile, and International Partners</u>
- Alerts to variable sources ("transient alerts") will be <u>available world-wide</u> within 60 seconds, using standard protocols
- LSST <u>data processing stack will be free software</u> (licensed under the GPL, v3-or-later)

All science will be done by the community (not the Project!), using LSST's data products.

LSST From a Scientist's Perspective

- A stream of ~10 million time-domain events per night, detected and transmitted to event distribution networks within 60 seconds of observation.
- A catalog of orbits for ~6 million bodies in the Solar System.
- A catalog of ~37 billion objects (20B galaxies, 17B stars), ~7 trillion observations ("sources"), and ~30 trillion measurements ("forced sources"), produced annually, accessible through online databases.
- Deep co-added images.
- Services and computing resources at the Data Access Centers to enable user-specified custom processing and analysis.
- Software and APIs enabling development of analysis codes.

Level 3



Center

French DAC

Satellite Processing

CC-IN2P3, Lyon, France

CONTRACTOR CONTRACTOR (50%)

Archive Site

NCSA Archive Center Alert Production Data Release Production (50%) Calibration Products Production EPO Infrastructure Long-term Storage (copy 2)

Data Access Center

Data Access and User Services



Dedicated Long Haul Networks

Two redundant 100 Gbit links from Santiago to Florida (existing fiber) Additional 100 Gbit link (spectrum on new fiber) from Santiago – Florida (Chile and US national links not shown)

Summit and Base Sites

Telescope and Camera Data Acquisition Crosstalk Correction Long-term storage (copy 1) Chilean Data Access Center

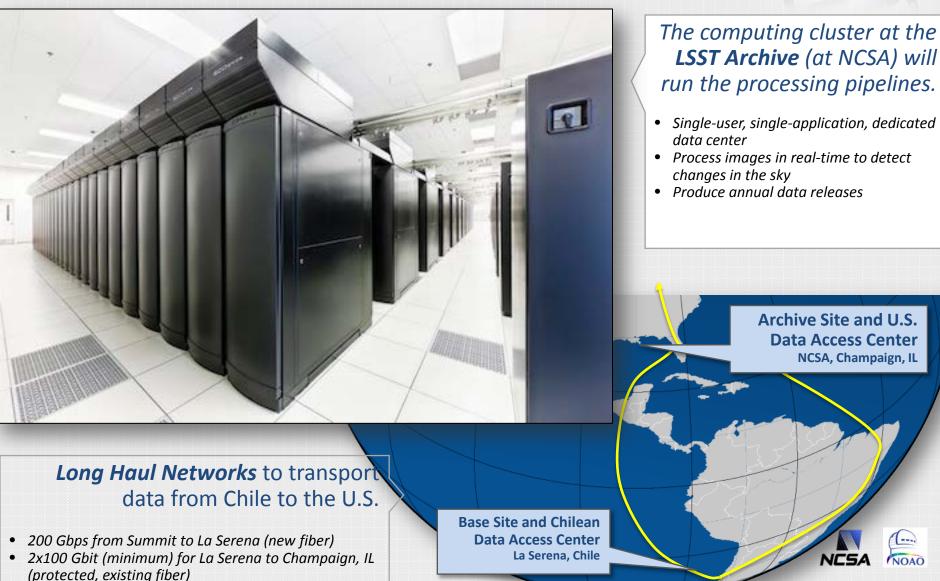


Science Operations Observatory Management Education and Public Outreach

HQ Site

Infrastructure: Petascale Computing, Gbit Networks





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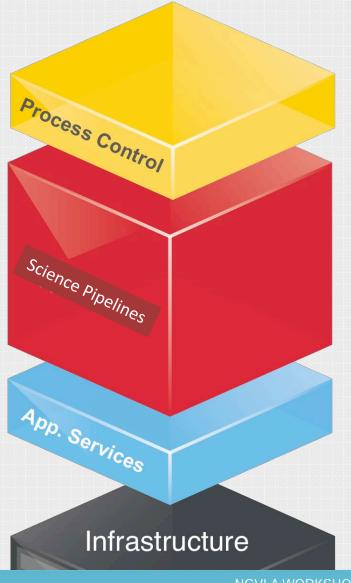
NESA

1----

NOAC

Science Pipelines: Scientific Core of LSST DM





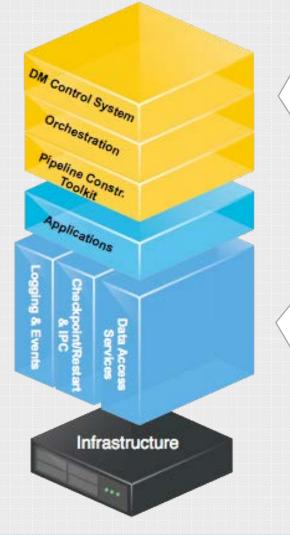
- Science Pipelines carry core scientific algorithms that process or analyze raw LSST data to generate output Data Products
- N.b. also referred to as "Applications"
- Variety of processing
 - Image processing
 - Measurement of source properties
 - Associating sources across space and time, e.g. for tracking solar system objects





Middleware Layer: Isolating Hardware, Orchestrating Software





Enabling execution of science pipelines on hundreds of thousands of cores.

- Frameworks to construct pipelines out of basic algorithmic components
- Orchestration of execution on thousands of cores
- Control and monitoring of the whole DM System



Isolating the science pipelines from details of underlying hardware

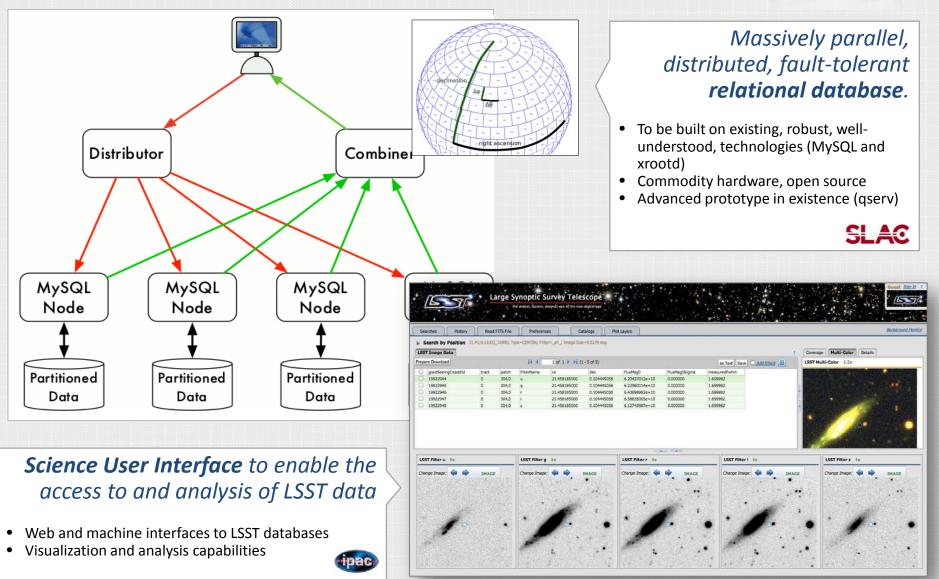
- Services used by applications to access/produce data and communicate
- "Common denominator" interfaces handle changing underlying technologies



SLAC

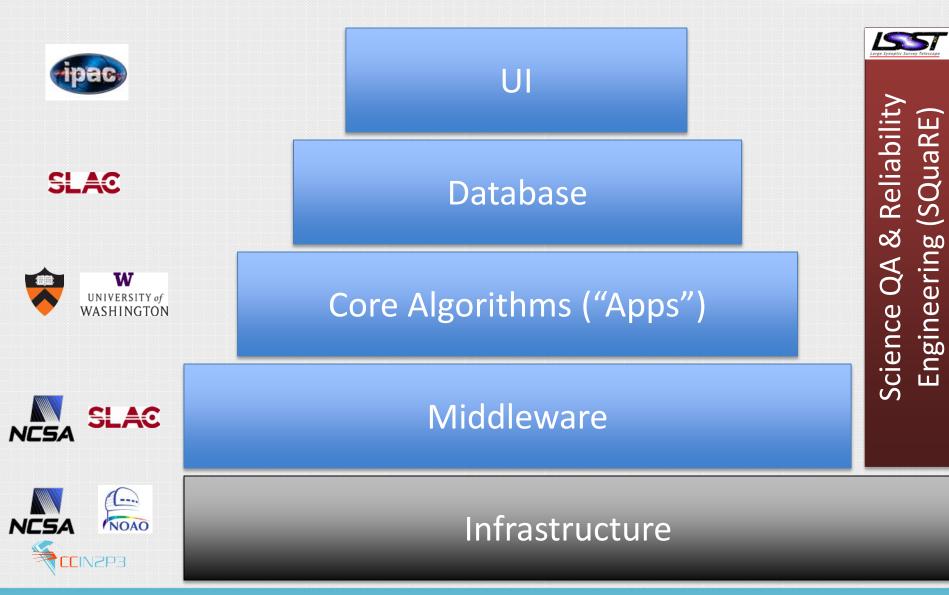
Database and Science UI: Delivering to Users





Distributed Team, SQuaREd





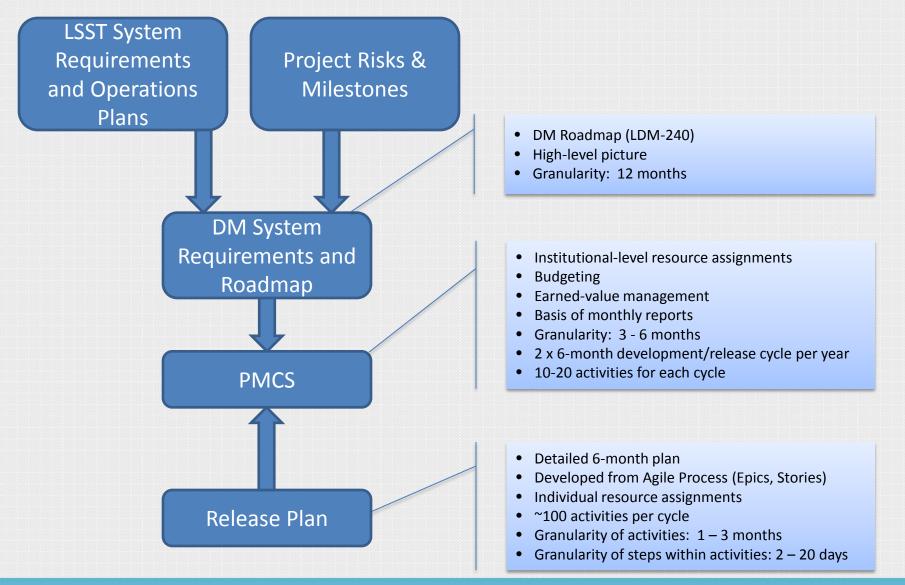
Tools and Processes



- Mailing lists, wikis, Confluence, telecons, videocons, ...
- Github
- Jenkins
- JIRA planning and Earned Value
- Licensing and contributions
- Investigating Slack and Discourse

Data Management Planning Process/Levels





LSST Software Development Roadmap (LDM-240

LSST Data I	Managen	nent Software	Development Ro	padmap				LDM-2 V.		
elease inish Date ear during which	Institution	R4.0 8/31/14	R5.0/5.1 8/31/15	R6.0/6.1 8/31/16	R7.0/7.1 8/31/17	R8.0/8.1 - #1 Release for ComCam 8/31/18	for ComCam 8/31/19	R10.0/10.1 - LSST Camera 8/31/20		
work is done		FY14	FY15	FY16	FY17	FY18	FY19	FY20		
			← Feature Implementation			← Performance Improvement				
			Source detection and measurement of PSF flux, aperture flux, adaptive moments	WCS terms needed to describe distortions at edges of chips and near the bloom stop						
2C.03.01 Single ame Processing pelines Washington		WCS determination module capable of fitting WCS to multiple CCDs at a time (in a single visit). Snap processing, CR detection and removal	Astrometric registration of a stack of images. Pixel-level intensity- dependent PSF correction Multi-CCD sky background detemination	Recognition and retention of trails (moving objects) in visit creation ISR functional with as-built CCD and raft characteristics (based on characterization of sensor/raft prototypes)						
			Improved sky background estimation Preliminary deblender	Cross-talk removal	Measureme functional with as CD and raft cb s.					
C.03.02 sociation seline	University of Washington	Source association and matchi refacto				Association of DIASources to DIAObjects Association of DIAObjects with Objects	Scalability and speed satisfy transient alert latency requirement			
C.03.03 Alert neration peline	University of Washington	MajoCells	or axes are 12 represent mission left are pr	ilestones/ma	jor results		Transient Alert Packet creation and transmission capability Initial implementation of simple Event Broker	Scalability and speed satis transient alert latency requirement Full implementation of simple Event Broker, abilit to receive and broker even under expected event and user load.		
		the		Source detection on difference images using	armacts (e.g., oright stars or ghosts)		Selection criteria and strategy for building templates Forced photometry on			

http://ls.st/ldm-240

LDM-240 DM Key Performance Metrics



Release	Unit	Contributing Elements	R4.0	R5.0/5.1	R6.0/6.1	R7.0/7.1	R8.0/8.1 - #1 Release for ComCam	R9.0/9.1 - #2 Release for ComCam	R10.0/10.1 - LSST Camera	Operational Readiness
lapsed time (years)			1	2	3	4	5	6	7	record in 1000
inish Date			8/31/2014	8/31/2015	8/31/2016	8/31/2017	8/31/2018	8/31/2019	8/31/2020	
ear during which work is done			FY14	FY15	FY16	FY17	FY18	FY19	FY20	
notometric repeatability										
procCalRep	mmag	SFM	15	13	10	5	5	4	3.5	3
PA1gri	mmag	SFM, CPP, PhotoCal	15	13	12	8	8	6	5.5	5
PA1uzy	mmag	SFM, CPP, PhotoCal	15	13	14	12	12	8	8	7.5
hotometric Spatial Uniformity										
PA3u	mmag	PhotoCal					40	30	20	20
PA3 (q)	mmag	PhotoCal					30	15		10
PA3 (y)	mmag	PhotoCal					30	15	10	10
Color Zero-point Accuracy										
PA5	mmag	PhotoCal					30	20	10	5
PA5u	mmag	PhotoCal			4		50	30	15	10
Absolute Photometry Accuracy										
PA6	mas	PhotoCal						30	20	10
Relative Astrometry		7100000				\prec			20	
AM1	mas	SFM, ObjChar	80	60 60			30 30	20		10
AM2	mas	SFM, ObjChar SFM, ObjChar, AstroCal	80	60			30	20	15	10
AM3 AB1	mas mas	SFM, ObjChar, AstroCal	100				4.0			
	mus	SPW, Objentil, Astrocal		• \	/laior a	ixes ar	e 12-m	onth cv	cles x r	metric
Absolute Astrometry										
AA1	mas	AstroCal		• ('alle rai	nracar	nt plann	ad achi	iovomo	nt of i
esidual PSF Ellipticity Correlations						preser	π μιαπη	ieu acin		
TE1	-	PsfEst, ObjChar, Coadd							r	.
TE2	-	PsfEst, ObjChar, Coadd		• P	rogres	sively	increas	ing per	formar	ice fro
loving Object Linkage Efficiency								01		
orbitCompleteness	%	MOPS		t	o right					
	<i>70</i>	MOr5		U						
puriousness Metric Efficiency								s in at f:		
transCompletenessMin	%	Diffim			RD MII	<u>nmum</u>	n spec is	<u>5 met TI</u>	rst, t <u>ne</u>	en de <u>si</u>
transPurityMin	%	Diffim								
mopsCompletenessMin	%	Diffim		+	hen (if	annlic	able) st	retch		
mopsPurityMin	%	Diffim		L C		applic				
				240	240	240	180	120		60
	seconds	All AP Pipelines + MW + Infra		231	193	154	77	58		39
omputational Performance Metrics OTT1 AP computational budget	TFLOPS	All AP Pipelines + MW				161	151	129		108
omputational Performance Metrics OTT1 AP computational budget DRP computational budget (DR1)	TFLOPS TFLOPS	All AP Pipelines + MW All DRP+AP Pipelines		645	215					
Computational Performance Metrics OTT1 AP computational budget DRP computational budget (DR1) LV query rate	TFLOPS TFLOPS simult. queries	All AP Pipelines + MW All DRP+AP Pipelines qserv		645 50	50	50	70	80		100
Computational Performance Metrics OTT1 AP computational budget DRP computational budget (DR1) LV query reponse time LV query reponse time	TFLOPS TFLOPS simult. queries seconds	All AP Pipelines + MW All DRP+AP Pipelines gserv gserv gserv		645 50 18	50 15	50 12	70 12	12	10	10
Computational Performance Metrics OTT1 AP computational budget DRP computational budget (DR1) LV query response time HV shared scan duration	TFLOPS TFLOPS simult. queries seconds hours	All AP Pipelines + MW All DRP+AP Pipelines qserv qserv qserv qserv		645 50 18 24	50 15 24	50 12 20	70 12 18	12	10	10
Computational Performance Metrics OTT1 AP computational budget DRP computational budget (DR1) LV query rate LV query response time HV shared scan duration HV simultaneous queries	TFLOPS TFLOPS simult. queries seconds	All AP Pipelines + MW All DRP+AP Pipelines gserv gserv gserv		645 50 18	50 15	50 12	70 12	12	10	10
omputational Performance Metrics OTT1 AP computational budget DRP computational budget (DR1) LV query rate LV query response time HV shared scan duration HV simultaneous queries	TFLOPS TFLOPS simult. queries seconds hours	All AP Pipelines + MW All DRP+AP Pipelines qserv qserv qserv qserv		645 50 18 24	50 15 24	50 12 20	70 12 18	12	10	10
omputational Performance Metrics OTT1 AP computational budget DRP computational budget (DR1) LV query rate LV query response time HV shared scan duration HV simultaneous queries	TFLOPS TFLOPS simult. queries seconds hours	All AP Pipelines + MW All DRP+AP Pipelines qserv qserv qserv qserv		645 50 18 24	50 15 24	50 12 20	70 12 18	12	10 12 12	10
Computational Performance Metrics OTT1 AP computational budget DRP computational budget (DR1) LV query response time HV shared scan duration HV simultaneous queries system Reliability Metrics	TFLOPS TFLOPS simult. queries seconds hours queries	All AP Pipelines + MW All DRP+AP Pipelines gserv gserv gserv gserv		645 50 18 24	50 15 24	50 12 20	70 12 18	12	10 12 12 5 0.5	1

1.02C.02 \$468,091 \$381,645 Performance Measurement 1.02C.02.01 \$102,727 \$102,727 **Baseline** (PMB) ▶ 1.02C.02.02 \$365,365 \$278,919 1.02C.03 \$339,445 \$151,181 2) Monthly export of Actuals is loaded 1.02C.03.00 \$64,192 \$60,027 into Cobra and is the basis for Actuals ▶<u>1.02C.03.01</u> \$31,882 \$31,882 (ACWP) NGVLA WORKSHOP - DECEMBER 8, 2015 - NRAO SOCORRO, NM



CV

(\$217,358)

(\$131,670)

(\$136.067

\$244,949

\$147,562

(\$71,320)

\$21,334

(\$111,285)

\$97,387

\$4,397

SV

(\$13,898)

\$229,508

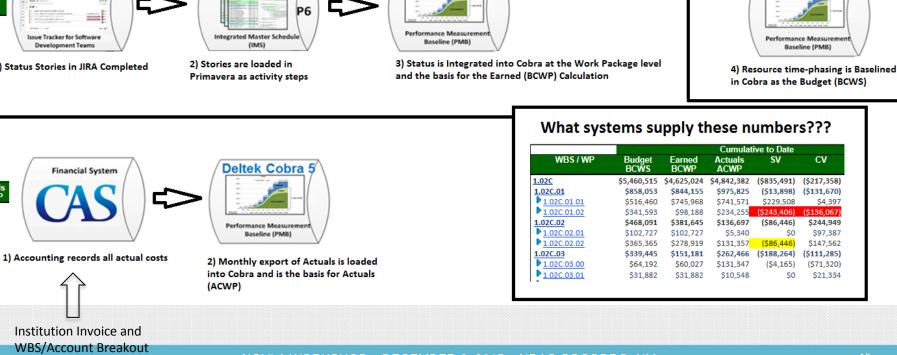
(\$86,446)

(\$86,446)

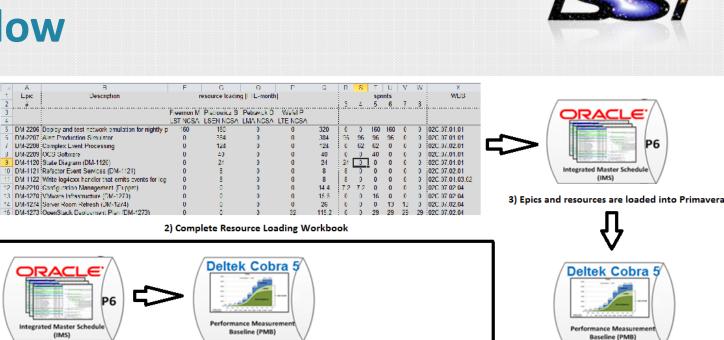
(\$4,165)

\$0

\$0



EV Data Flow



1) Status Stories in JIRA Completed

ŸJIRA

Issue Tracker for Software

Development Teams

1) Create Epics In JIRA

ŸJIRA

Budget BCWS

Earned BCWP

Actuals ACWP

Roadmap



- 2016: Finish end-to-end, start refining components, developer multipliers, run easily on precursor data
- 2017: Reprocess DES, continuous simulated Alert Production
- 2018: Ready for ComCam
- 2019: Ready for full Camera
- 2020: Ready for Science Verification

Team Culture (emphasis added,

truncated)



The LSST DM community includes LSST-paid staff at multiple institutions and external contributors from around the world. We have built a strong team that collaborates well. As we hire new people, we want to make sure we are maintaining a healthy, supportive, productive culture. While culture is best transmitted by daily example, having some formal standards for conduct can aid newcomers and reinforce good patterns.

Since LSST is an AURA center, the **AURA standards of workplace conduct** provide a starting point. As stated there, we in DM dedicate ourselves to fostering a **civil and inclusive community characterized by mutual respect for the contributions of all individuals**. As a community, we embrace the values in that document, in addition to any **local institutional standards**.

Lessons Learned



- AURA/LSST is in many ways a completely new entity, with inherent start-up activities
- Hiring has been slow, particularly for lead roles
- Team is not yet developing at maximum productivity
- It is hard to mesh scientists, managers, and engineers on one team
- Algorithmic progress is hard to measure and even harder to predict
- Architectures are evolving rapidly: e.g. Cloud, AstroPy, Hadoop ecosystem
- It is hard to balance technical priorities, trading off between performance/scale versus adaptability and usability
- Reorganize work with delayed staffing, relieve current leads of less-critical tasks
- Focus on developer multipliers
- Establish team culture and norms
- Focus on "end-to-end" system
- Document design path and replan
- Track new architectures, maintain portability
- Make sure we don't accumulate too much "technical debt"