

Design and Implementation of the ACSIS Real-Time Data Reduction Pipeline

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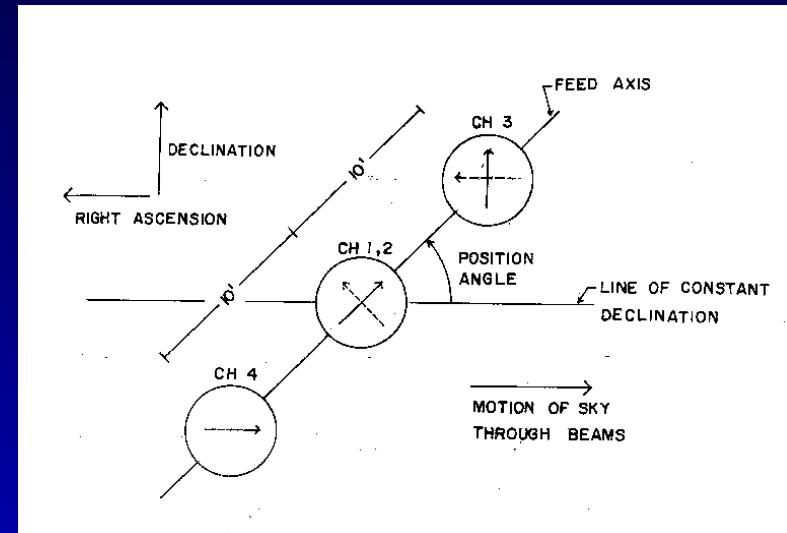
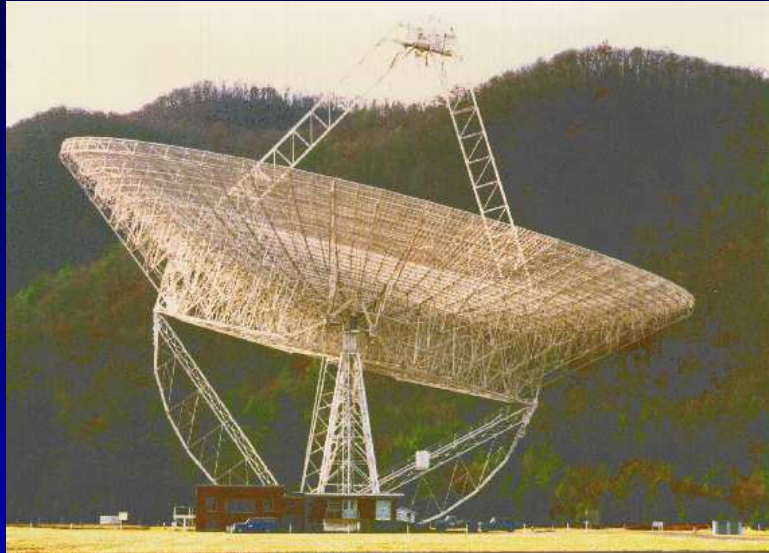
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History Repeats Itself

- 300 ft telescope with 3-feed focal plane array at 11 cm wavelength (1971)



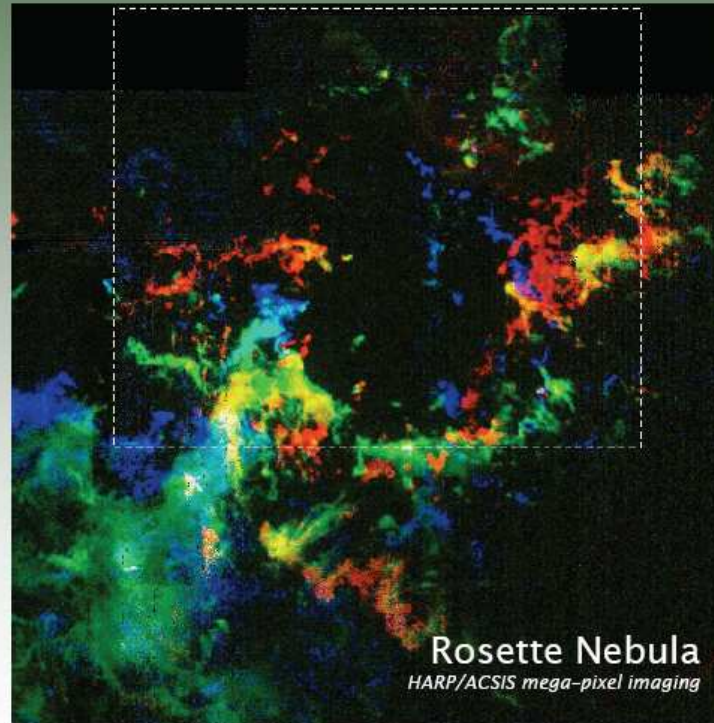
ACSIS Does Work!



JCMT SPECTRUM

NEWSLETTER OF THE JAMES CLERK MAXWELL TELESCOPE

SPRING 2007 · #26

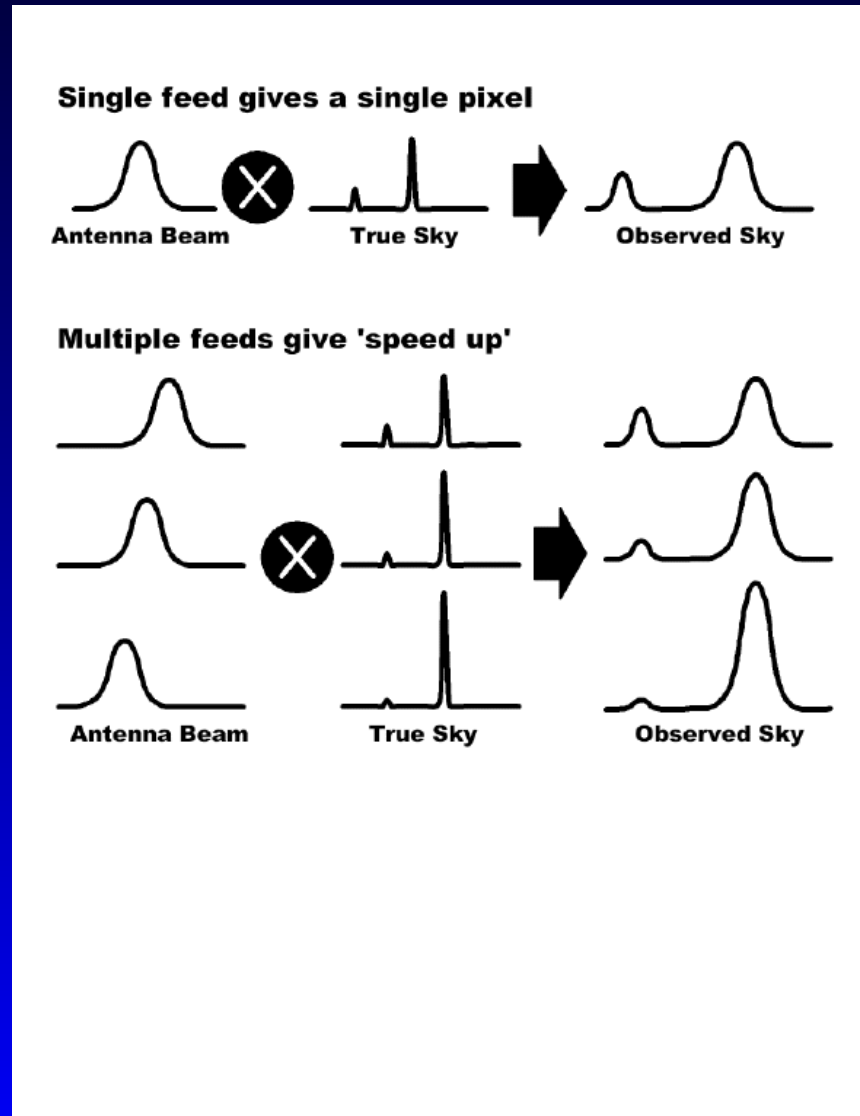


Rosette Nebula
HARP/ACSIS mega-pixel imaging

Early HARP Results

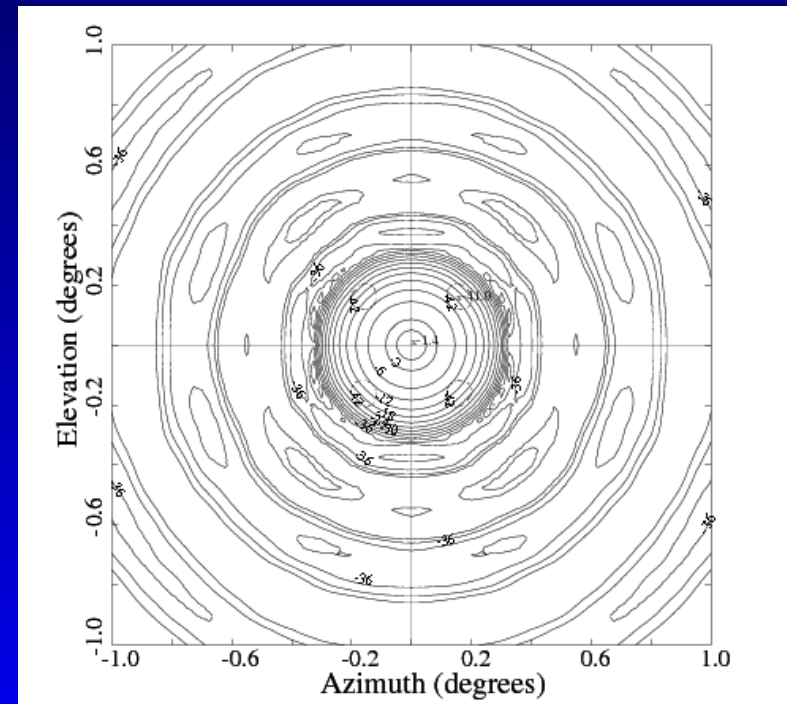
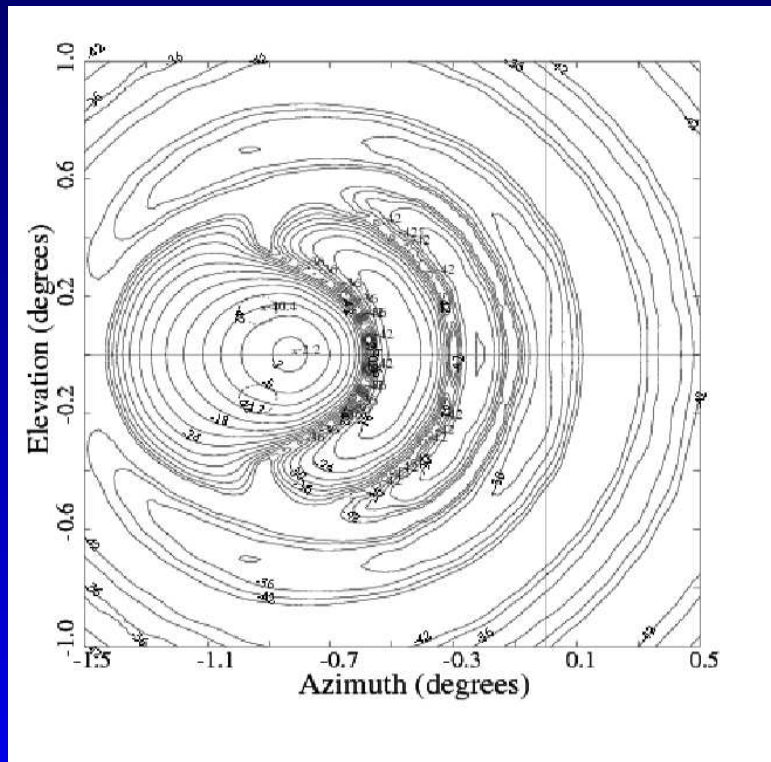
Why We're Here

- Focal plane arrays give enormous observation speedup



But No Free Lunch

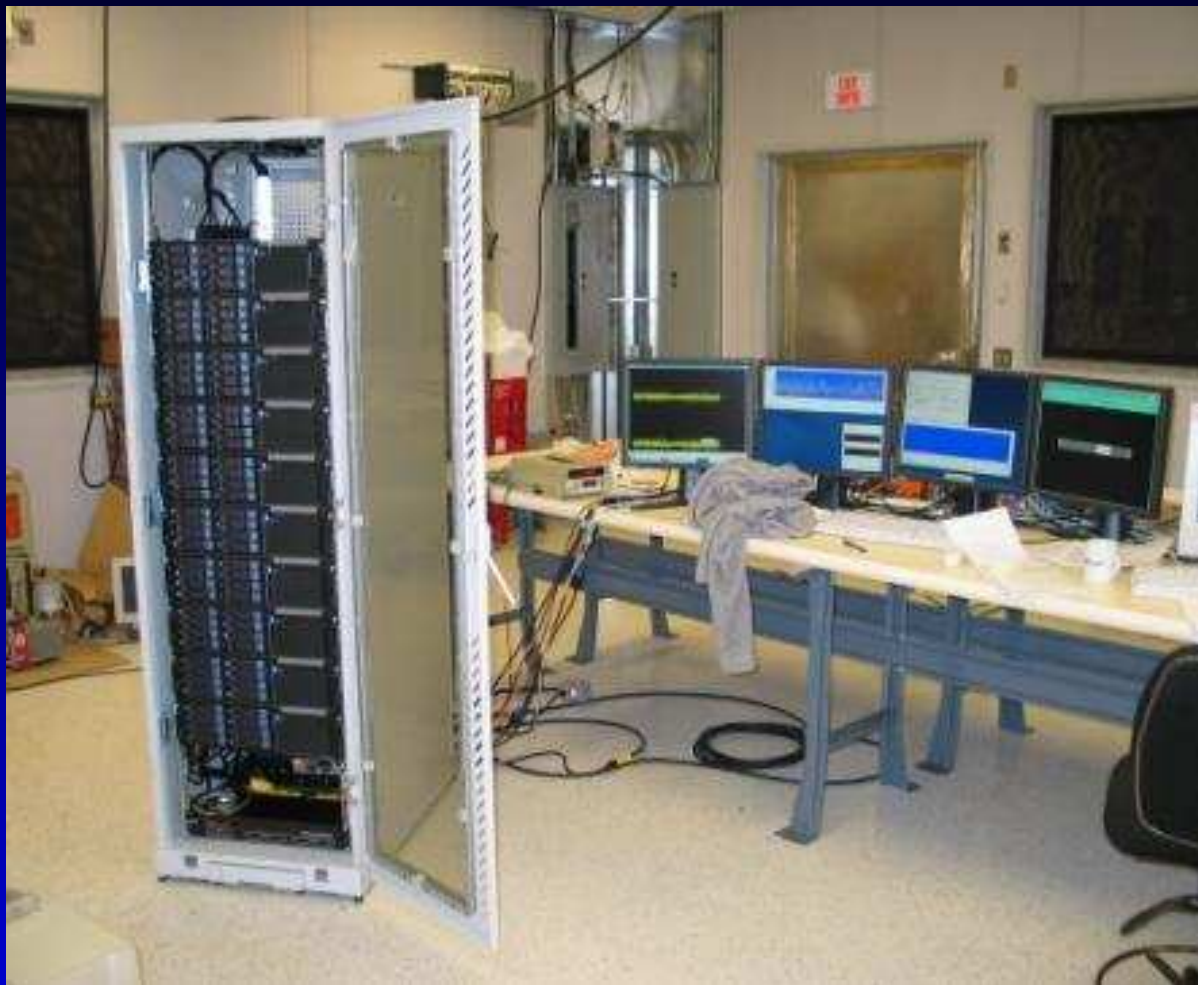
- Off-axis beams can have ugly shape and high instrumental polarization
- The figures show the central and one of the circumferential beams of the Parkes 21 cm multi beam system



Main ACSIS Configuration

- HARP receiver has 16 pixels
- 16×8192 channel mode with 50ms integration time produces 10 Mb of data per second
- originally we planned a distributed system for data processing on a small ‘beowulf’ cluster of 32 450MHz PCs
- final system had 10 dual processor 2GHz PCs
- parallel system needed because of
 - data rate vs available CPU speed
 - 32-bit memory limitations (image cube has to be spread over multiple machines)

ACSIS Computing Hardware



ACSIS Implementation

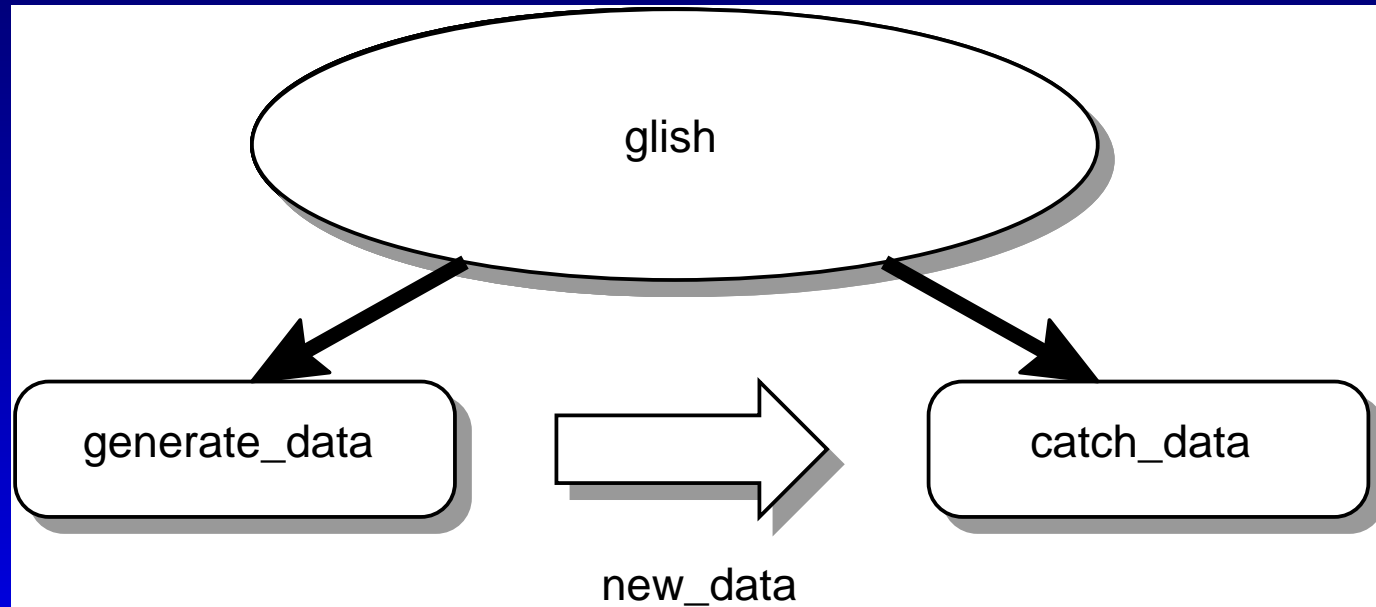
- use c++ class library provided by aips++
- client c++ programs are connected by glish scripting and messaging system
- communication takes place by means of events. Clients can send events directly to each other by 'links'
- glish is a 'good thing' for this type of system
- see aips++ Note 230 'Advanced Programming with AIPS++ and Glish' for a detailed discussion of distributed programming with glish
- system comes with simulation tasks for testing and debugging purposes

glush communication with client

```
##### begin hello.g #####  
# a glush record containing the  
# string 'hello world'  
info := [=]  
info.data_string := 'hello world'  
  
# start c++ glush client on machine slave5  
demo:= client("hello_world",host ="slave5")  
  
# send this client the glush 'info' record  
# by means of a hello event  
demo->hello(info)  
##### end hello.g #####
```

glish link

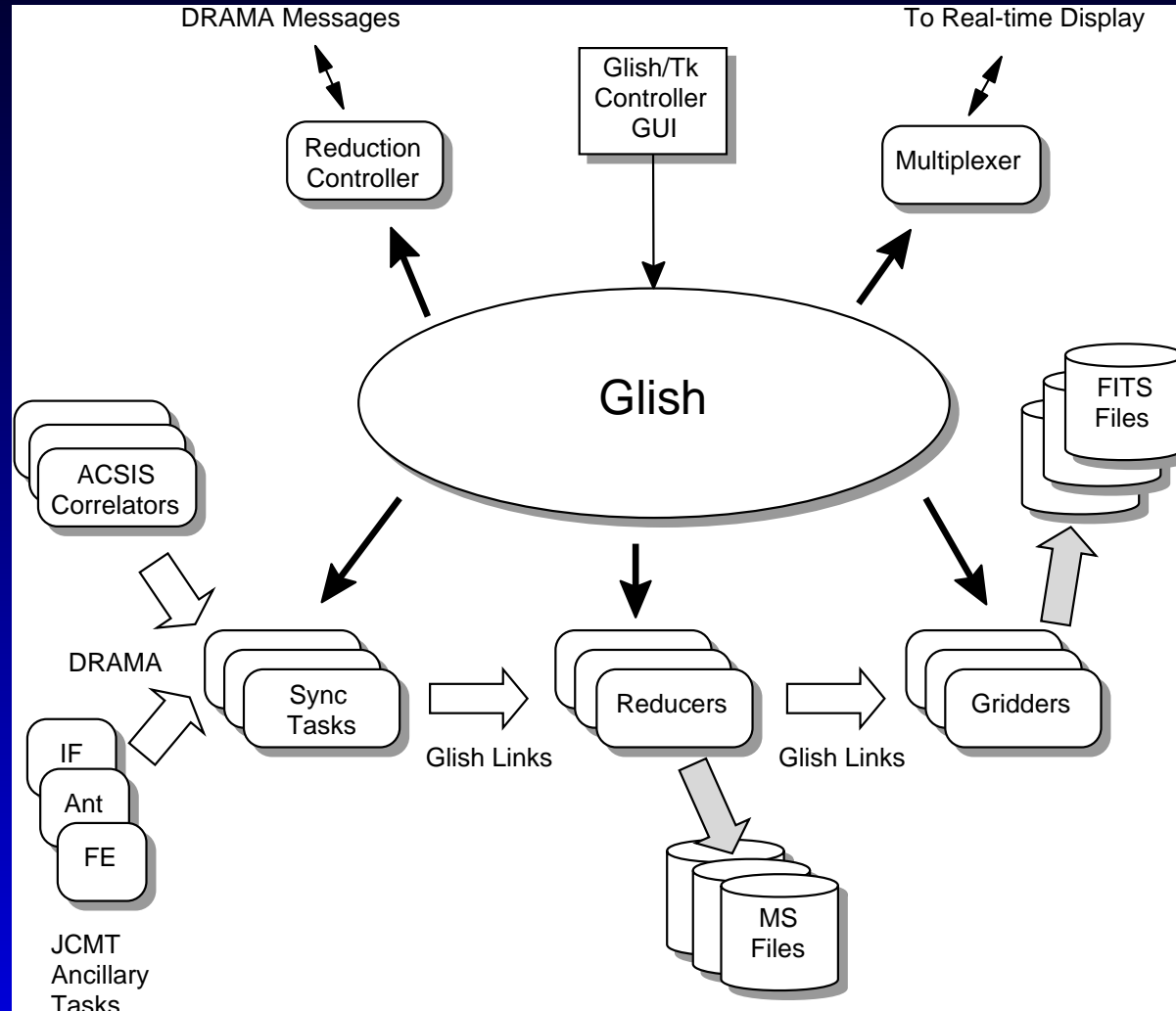
```
##### begin link.g #####  
a:= client("generate_data", host="slave1")  
b:= client("catch_data", host="slave2")  
link a->new_data to b->new_data  
##### end link.g #####
```



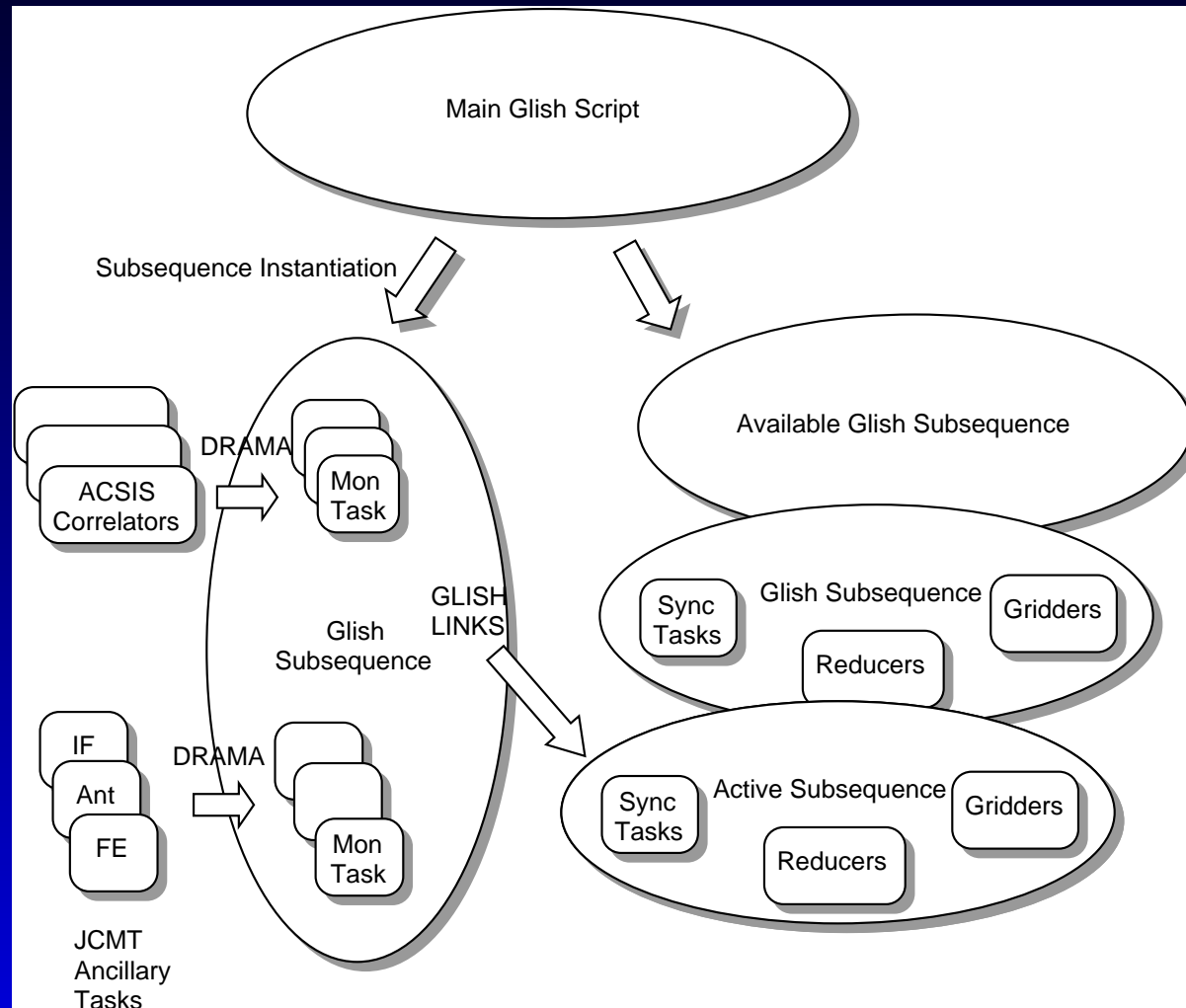
System Scalability?

- Yes!
- System can be configured to run on anything from a laptop to (at least) a 32 processor PC cluster
- Upper limit is unknown
- Perl scripts read XML-based configuration files to load and link tasks together across the network
- Can be used with any feed configuration from 1 to n where n has a maximum of ??

Reduction System Overview



Pipeline Details



Main ACSIS Tasks

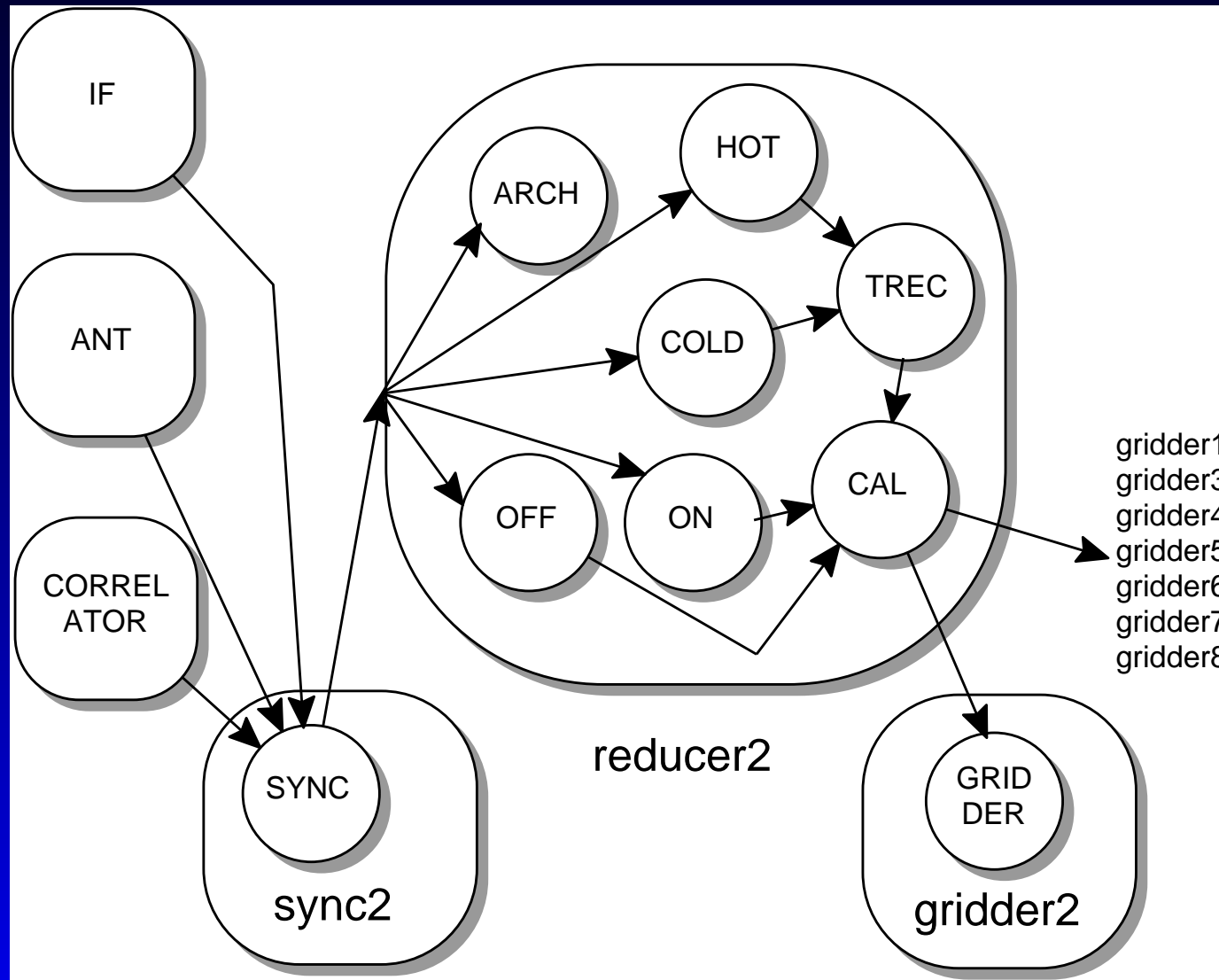
- Reduction controller: interface between the reduction system and the rest of the JCMT telescope control software
- Multiplexer: a shared glish client that can be connected to by multiple glish interpreters
- Sync task: synchronize data from a correlator crate with the ancillary data that are required to completely process a spectrum
- ReducerProcess: convert raw correlator lags (time domain) into calibrated spectra (frequency domain)
- Gridder: insert calibrated spectra into correct location of a 3-dimensional data cube

XML for Task Instantiation

- Glish scripts are essentially 'fixed'.
- System configures itself for each individual observation by reading and processing XML files.

```
<process-init>  
  <process name="sync2" exe="sync_task"  
                                         machine="slave2">  
  
  </process>  
  ...  
</process-init>
```

Reducer Process Instantiation



ReducerProcess XML

```
<rt-process type="REDUCER">  
  <red-object name="ON" type="reducer">  
    <red-object-source-list>  
      <string> SELECT FROM SYNC  
        WHERE LOAD=="SKY" && ON_SOURCE</string>  
    </red-object-source-list>  
    ...
```

XML for ReducerProcess con't

```
<red-object-recipe>
  <string>      corr_data.FLOAT_DATA </string>
  <function> rvalue </function>
  <string>      SPECTRAL_WINDOW_ID </string>
  <function> rvalue </function>
  <function> window </function>
  <function> fft </function>
  <string>      corr_data.FLOAT_DATA </string>
  <function> lvalue </function>
</red-object-recipe>
</red-object>
...
</rt-process>
```

RPN Algorithm Equivalent

- push incoming data field `corr_data.FLOAT_DATA` on to the stack
- pop data from the stack, multiply it by an appropriate window function and push result
- pop data from the stack, FFT it and store it in the `corr_data.FLOAT_DATA` field
- forward the updated data record to other objects that have ON in their `<red-object-source-list>` definitions

Flexibility

- Reducer Processes are basically programmable RPN calculators for 8000 channel vectors
- Can write a new program for a new observing procedure
- System was delivered to the JCMT with six ‘standard’ observing recipes

System Output

- Reducer Processes write raw input data to disk in the form of aips++ Measurement Sets
- Gridders write out (sub) data cubes as FITS files
- Data can be reprocessed with a 'replay' script. This script can read in the saved MS data and re-reduce it with a modified recipe (e.g. with different baseline fitting regions or a different gridding kernel)

Real Time Display Overview

- Real time displays available for
 - meta data
 - Reducer Process ‘nodes’
 - Gridders
- For meta data and Reducers, data are shown collectively, but can be examined on an individual basis
- meta data displays can run continuously
- new Reducer and Gridder displays are created for each observation
 - are created by XML specification
- displays provide immediate feedback during an observation

Data Selection GUIs

- GUI used to select data from either Reducer Processes (spectrum selector) or Gridders (image selector)
- Reducer Process ‘nodes’ can be examined individually (especially useful for debugging)
- GUIs and displays mostly use Qt widget library with QWT plotting widgets

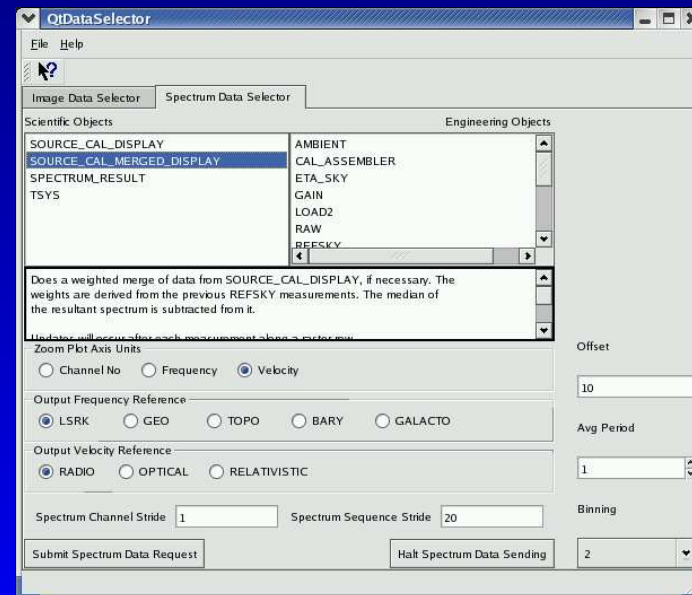
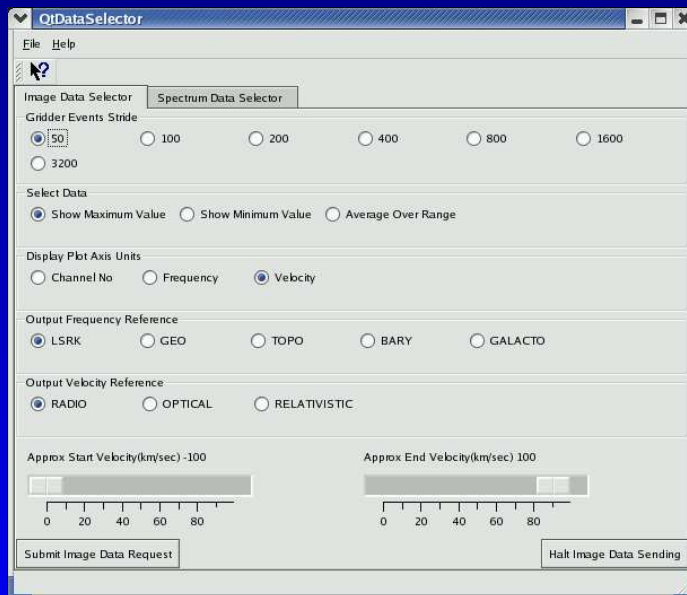
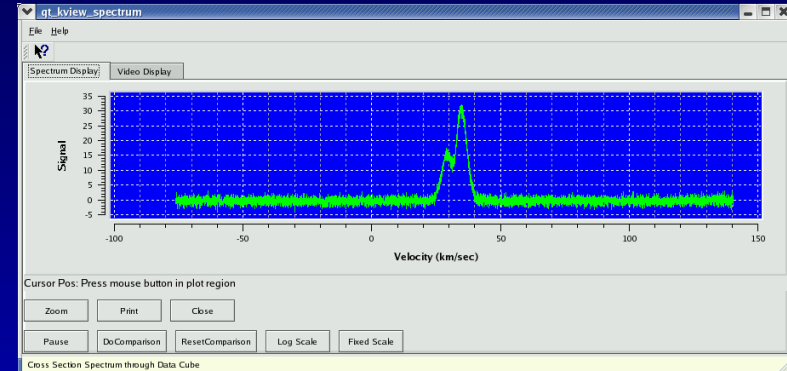
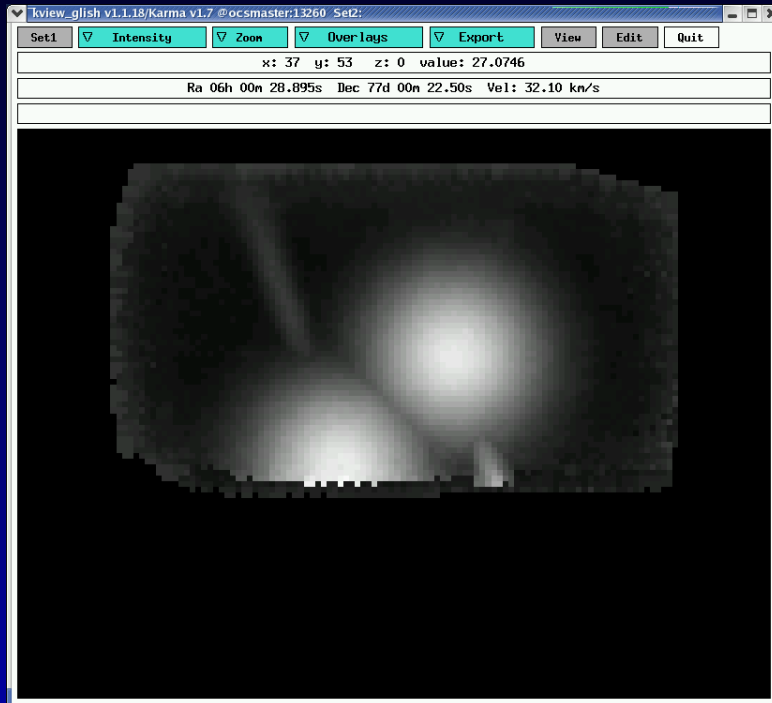
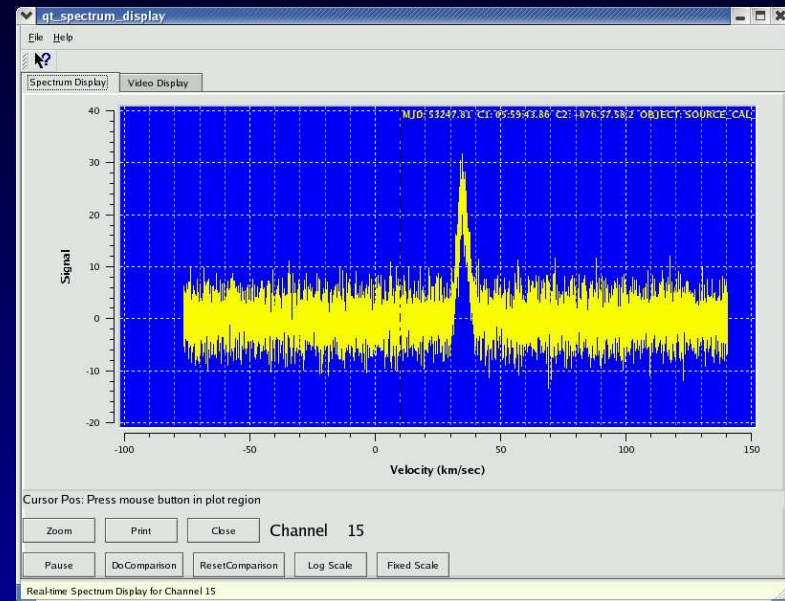
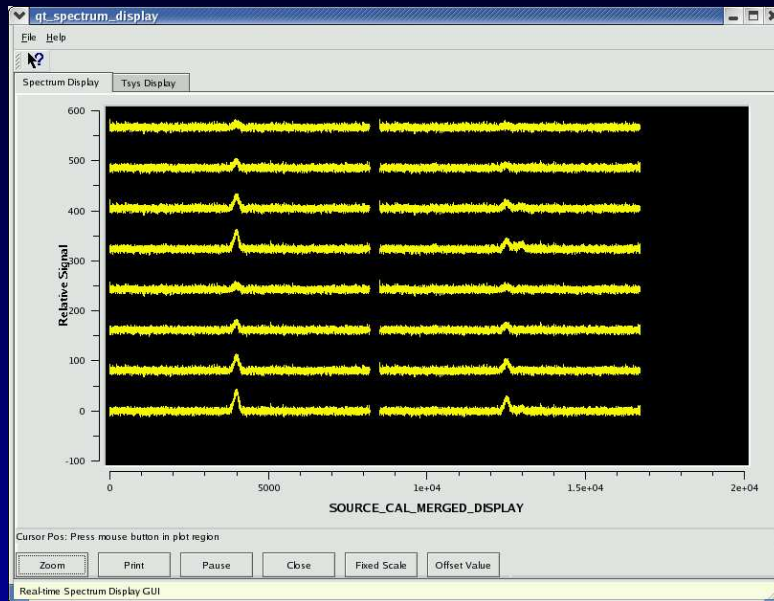


Image Displays



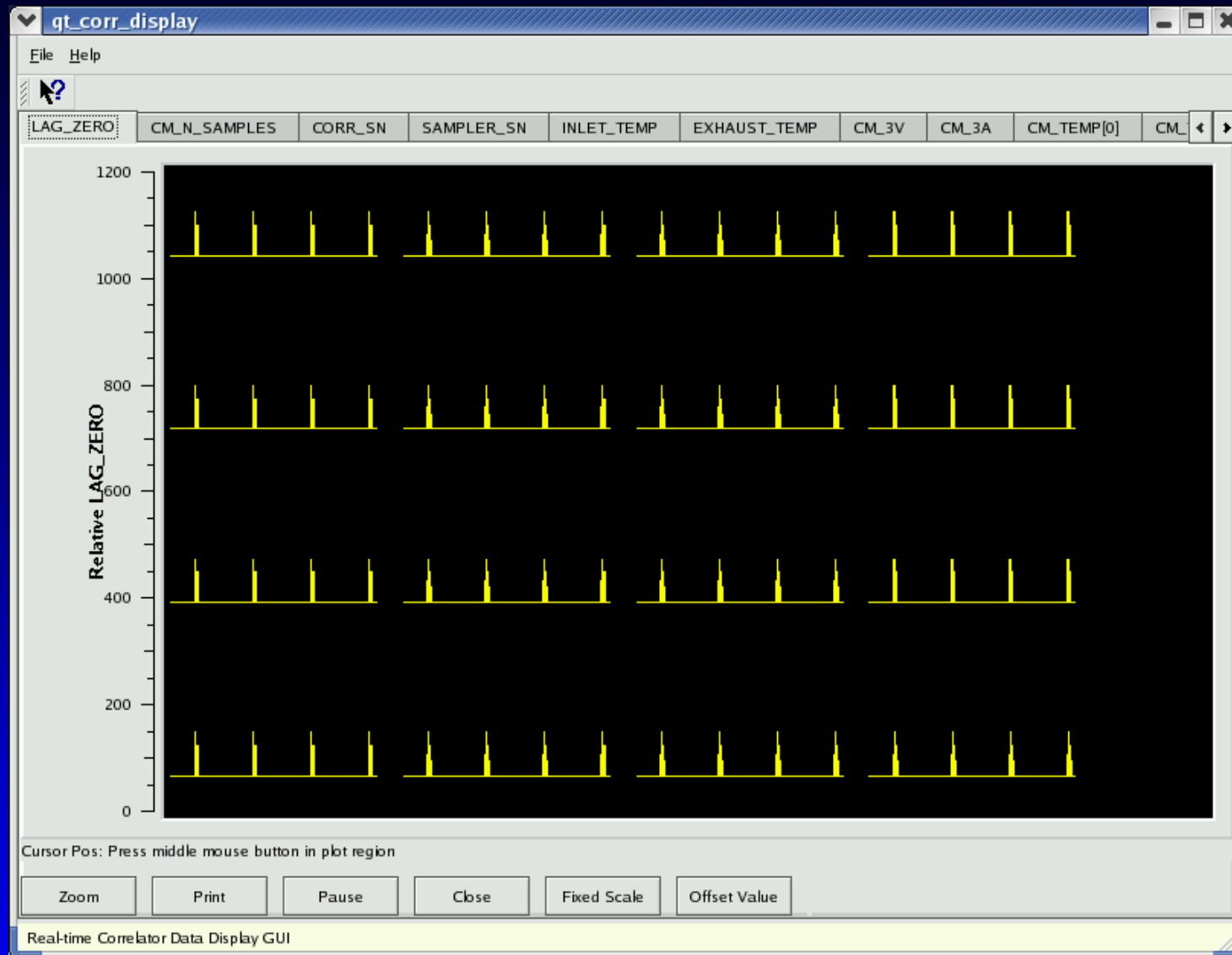
- left GUI shows kview continuously updating during an OTF imaging observation
- right GUI shows spectrum through the entire image cube at current position of the mouse on kview

Spectrum Displays



- left GUI shows all current spectra
- right GUI shows popup that appears when user clicks on an individual spectrum on the left GUI

Sample Meta data Display



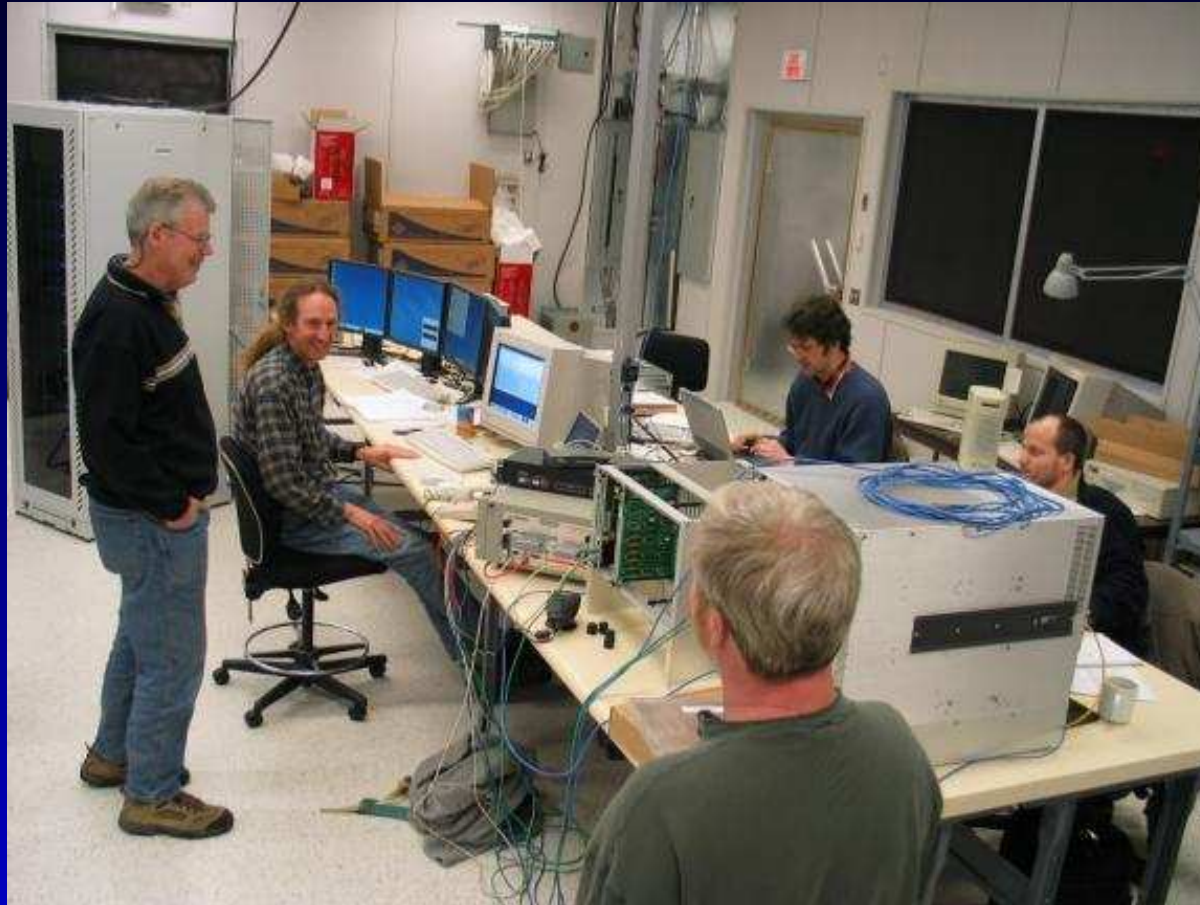
ACSIS Pipeline Summary

- Scalable
- Flexible
- Real time feedback on data quality
- Can view intermediate products for debugging and test purposes
- Outputs are aips++ MS for raw data and FITS files for image cubes
- MS data can be reprocessed (non-real time)
- Polarization capability could probably be added with a new reduction recipe

Lessons Learned

- A system for interprocess communication that provides approximately 400 different function calls is not necessarily better than a system that has only two function calls.
 - KISS (not the band)
- Inter agency planning and cooperation needed
 - JAC/DRAO/UKATC squabbles over resource availability, deliverables, delivery times etc
 - distributed development can work but it requires the right people

Some ACSIS People



- Tony Willis, Bill Dent, Gary Hovey, John Lightfoot and Tim Jenness during lab acceptance testing

Addendum

- Glish event loop successfully combined with
 - DRAMA
 - Qt widget set
 - CORBA - Orbacus
 - karma - kview (R. Gooch)
 - Xt (aips++ group)
 - MPI (indirectly - by aips++ group)
 - WWW Boyd Waters et al (ADASS 2002 paper)
- So it's quite possible to combine glish / aips++ with other messaging systems